

# High Contrast Imaging of Massive Stars

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San Carlos de Bariloche, Argentina



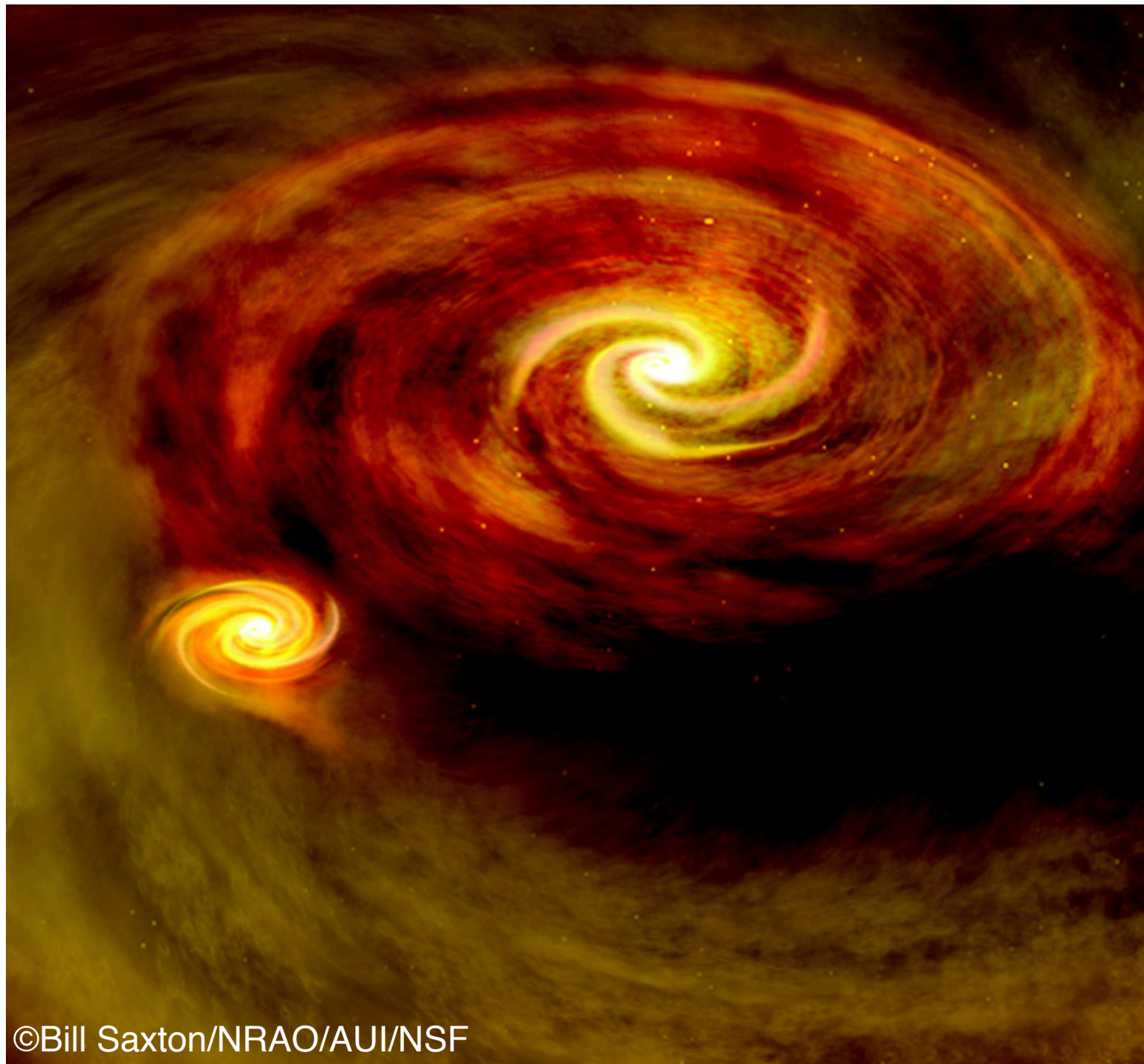
<https://arainot.github.io>

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**KU LEUVEN**

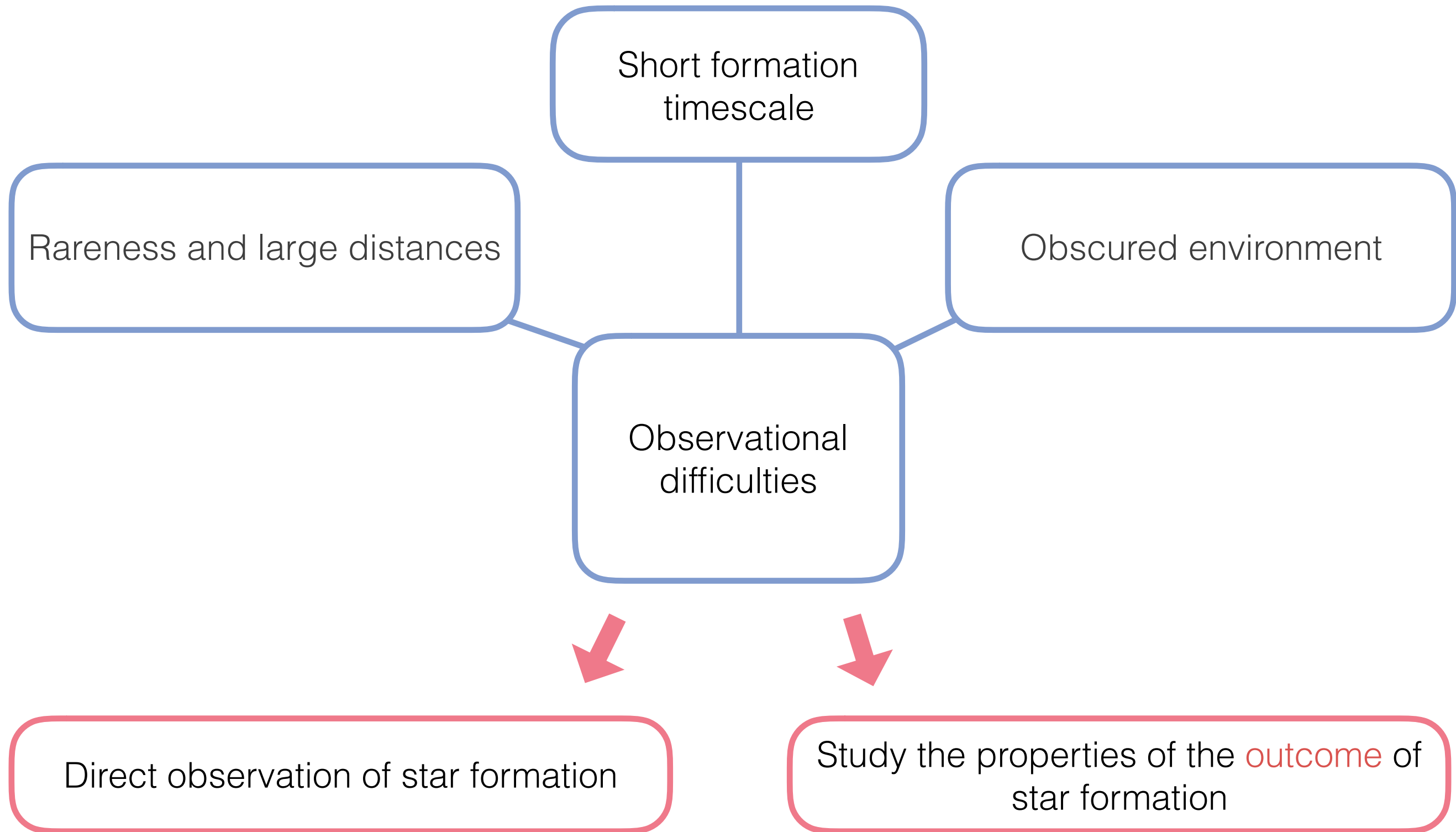
What are massive stars?

# Massive Star Formation



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# Massive Star Formation Problems





**C**

**H**

**I**

**P**

**S**

**C**arina

**H**igh contrast

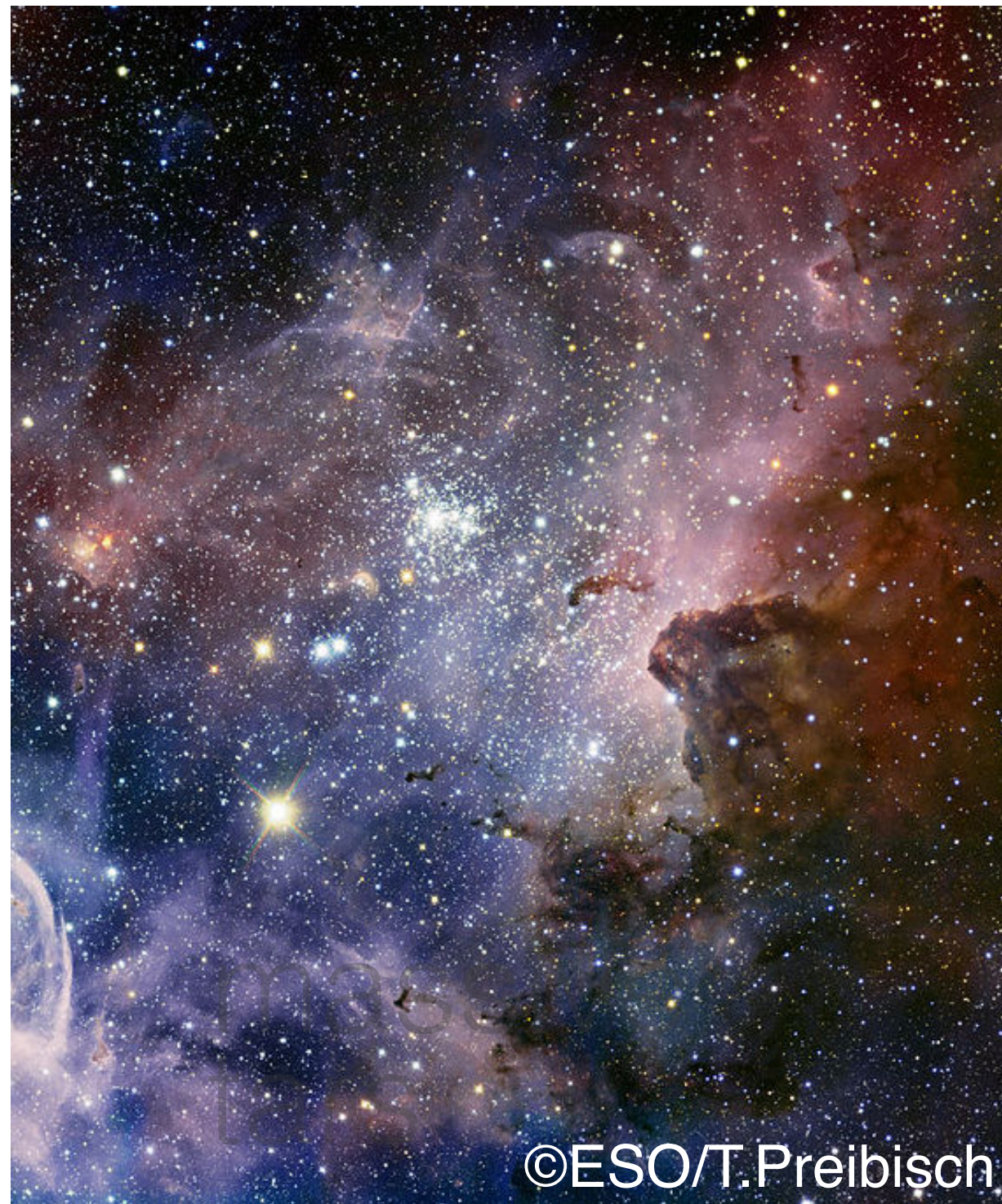
**I**maging

**P**roject of

**S**massive  
tars

# Carina

close massive star  
region



©ESO/T.Preibisch

**C**arina — close massive star region

**H**igh contrast  
**I**maging — VLT/SPHERE in IRDIFS mode

**P**roject of  
**S**massive stars



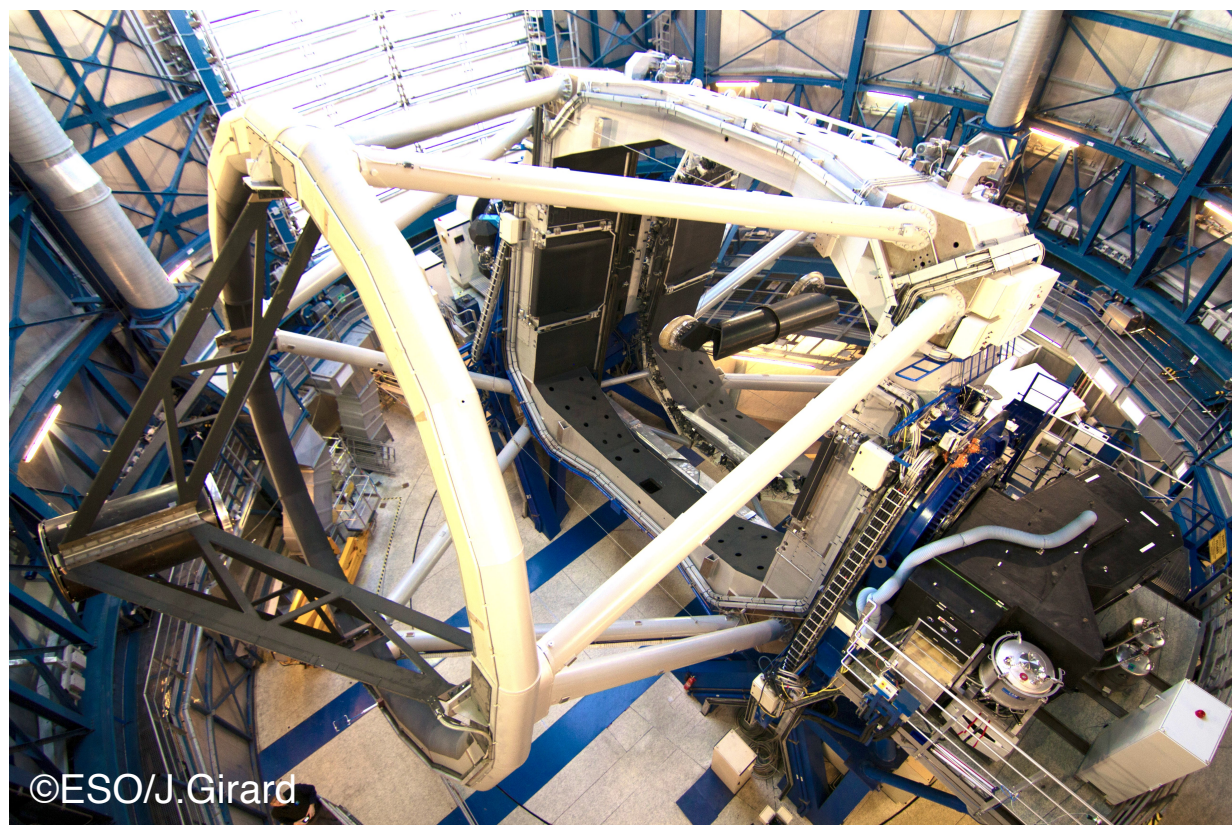
# Carina

close massive star region

# High contrast

VLT/SPHERE in IRDIFS mode

# Imaging



©ESO/J. Girard

SPHERE	IFS	IRDIS
<b>Spectral Range (<math>\mu\text{m}</math>)</b>	0.95-1.75	0.95-2.32
<b>FOV (<math>\text{arcsec}^2</math>)</b>	1.73	11
<b>Pixel Scale (marcsec)</b>	7.4	12.25
<b>Bands</b>	Y-J-H	K (1&2)

**C**arina — close massive star region

**H**igh contrast  
**I**maging — VLT/SPHERE in IRDIFS mode

**P**roject of — faint and low-mass companions

**S**massive stars

**C**arina

close massive star

63 stars in P102

**H**igh contrast

28 stars already reduced

VLT/SPHERE in IRDIFS mode

**I**maging

**P**roject of

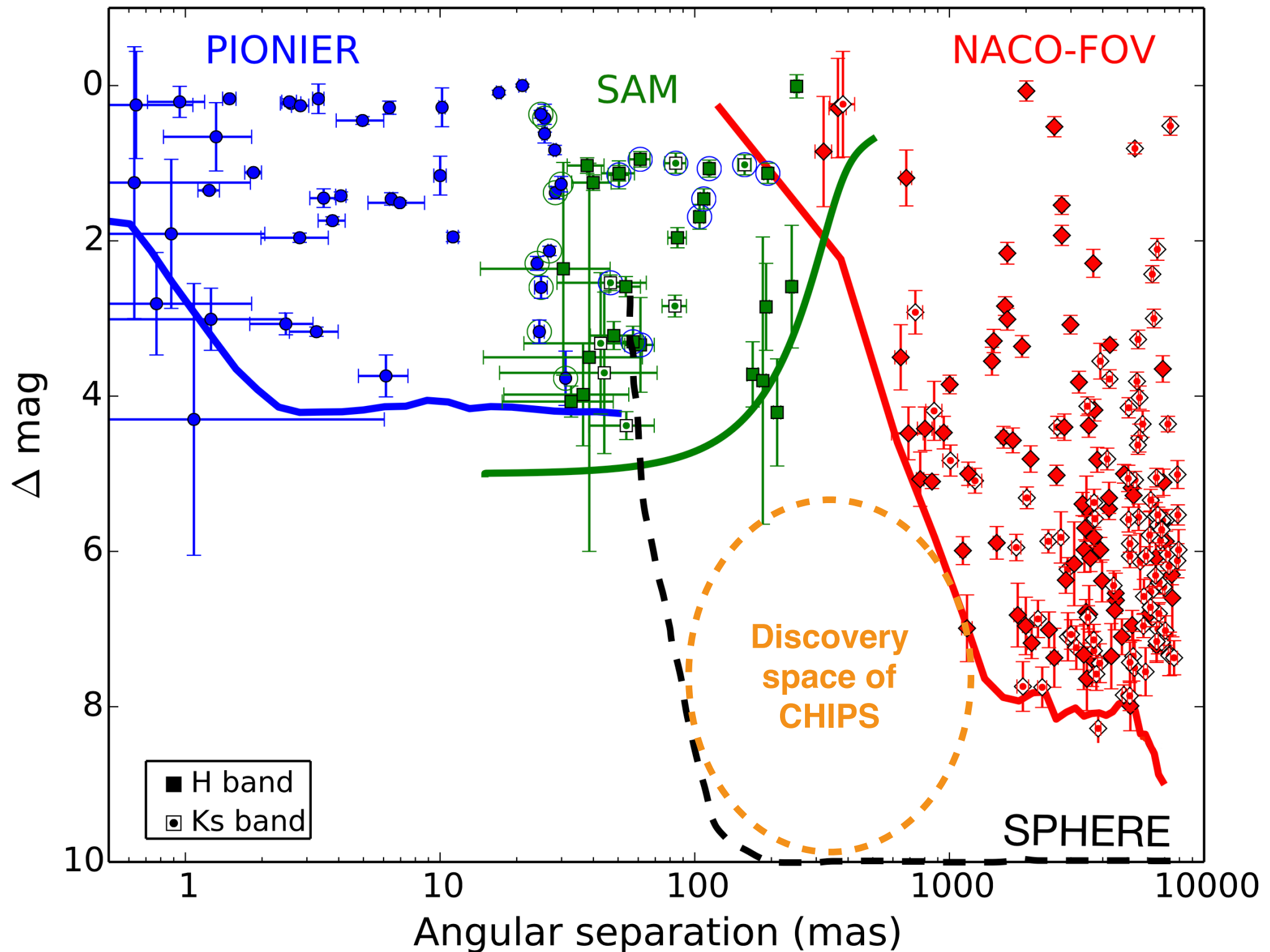
4D data cubes

**S** massive stars

Multiplicity properties of 91 massive O and WR



# SMaSH+ (Sana et al, 2014)

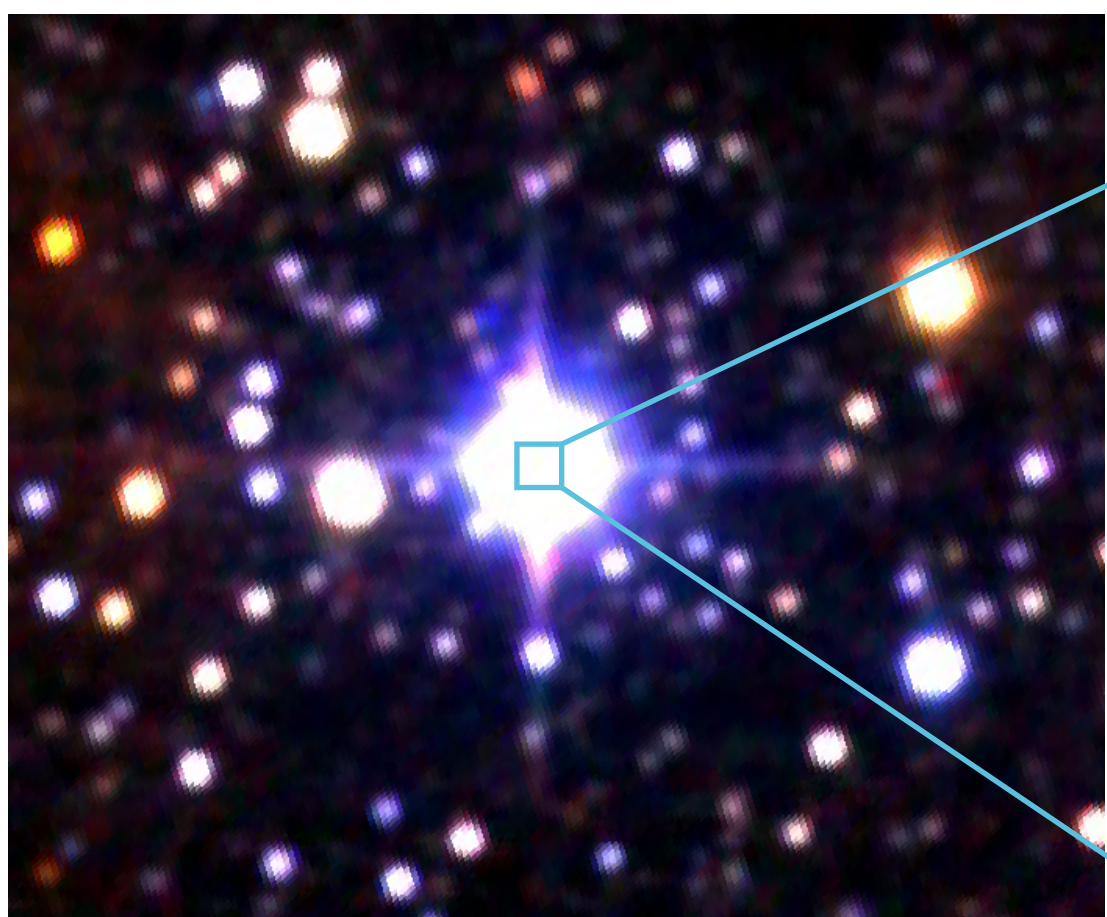


Some example data

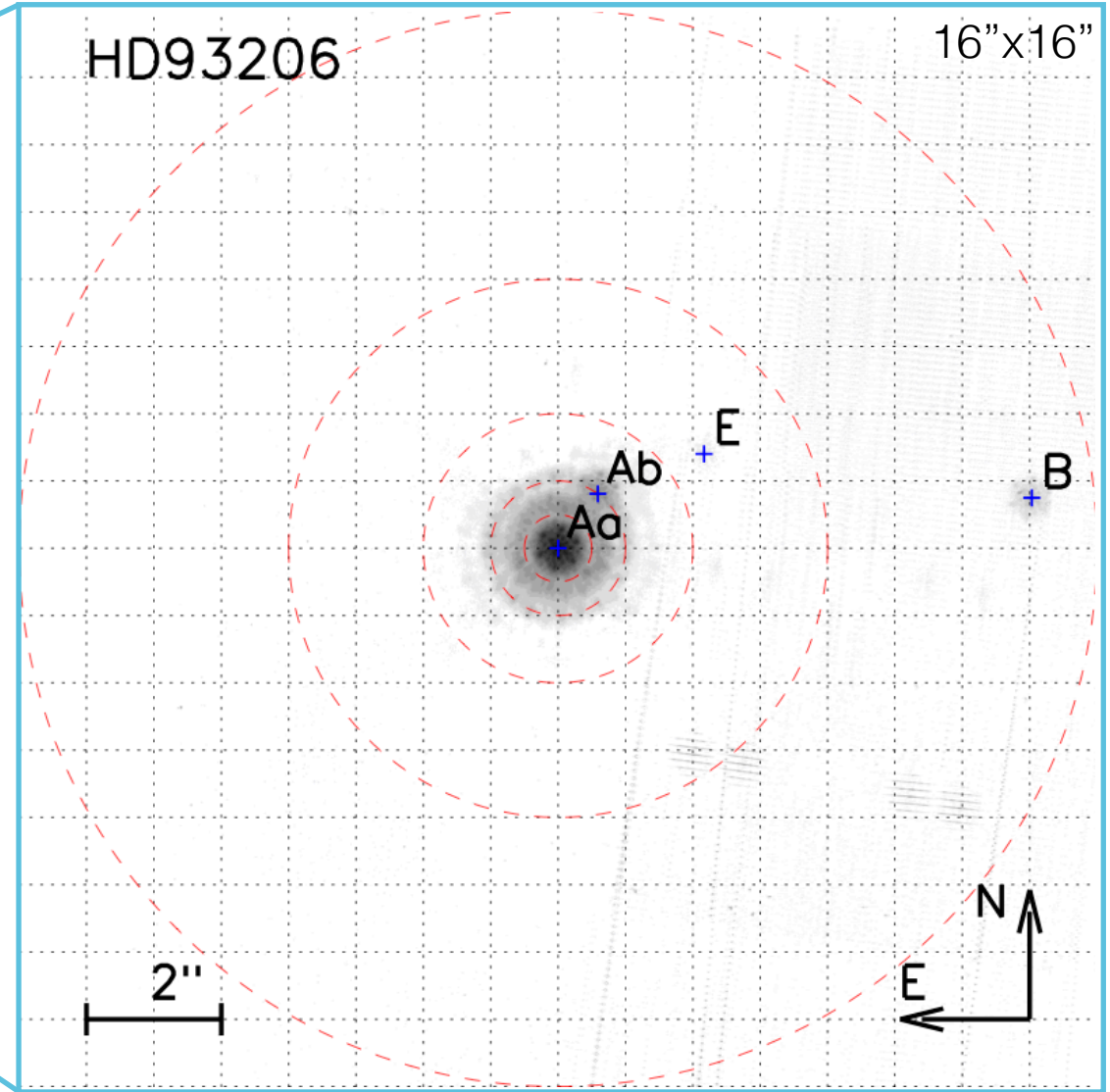
QZ Car (HD 93206)



**FOV = 5', 2MASS**

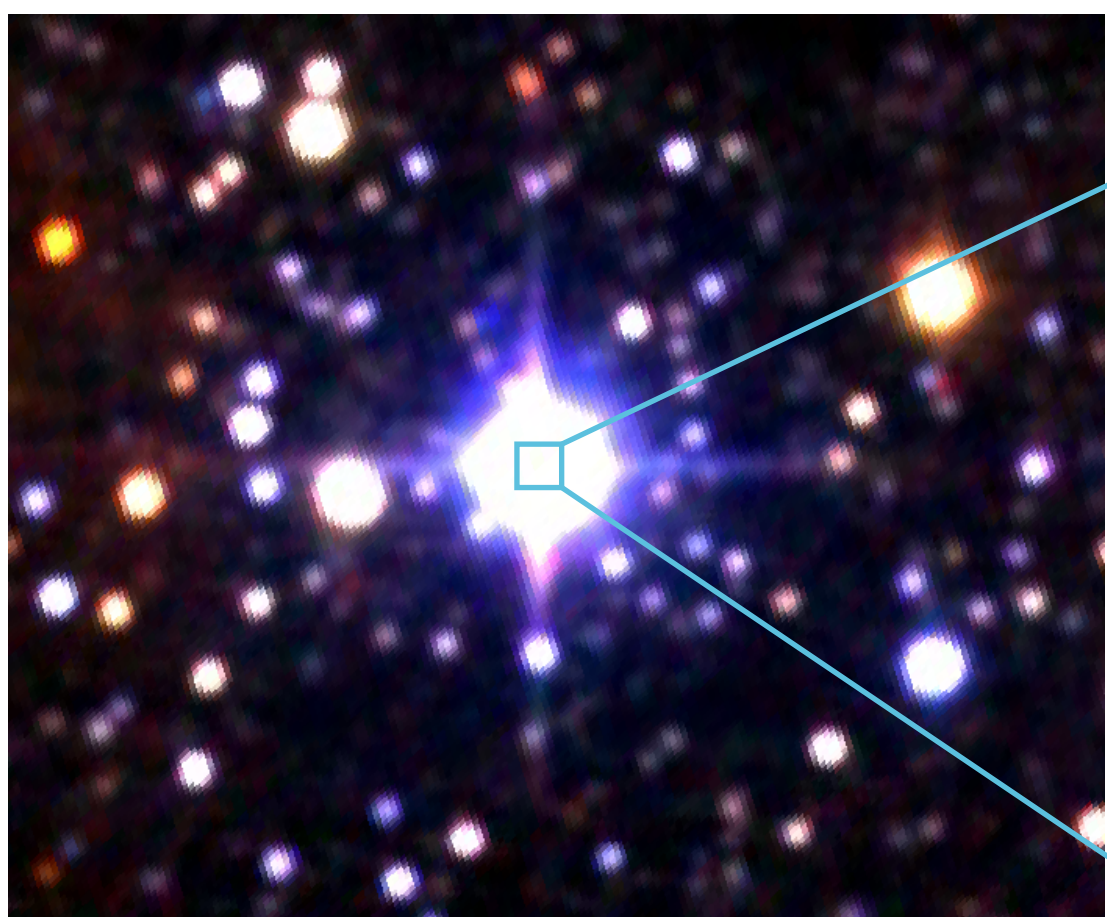


FOV = 5', 2MASS

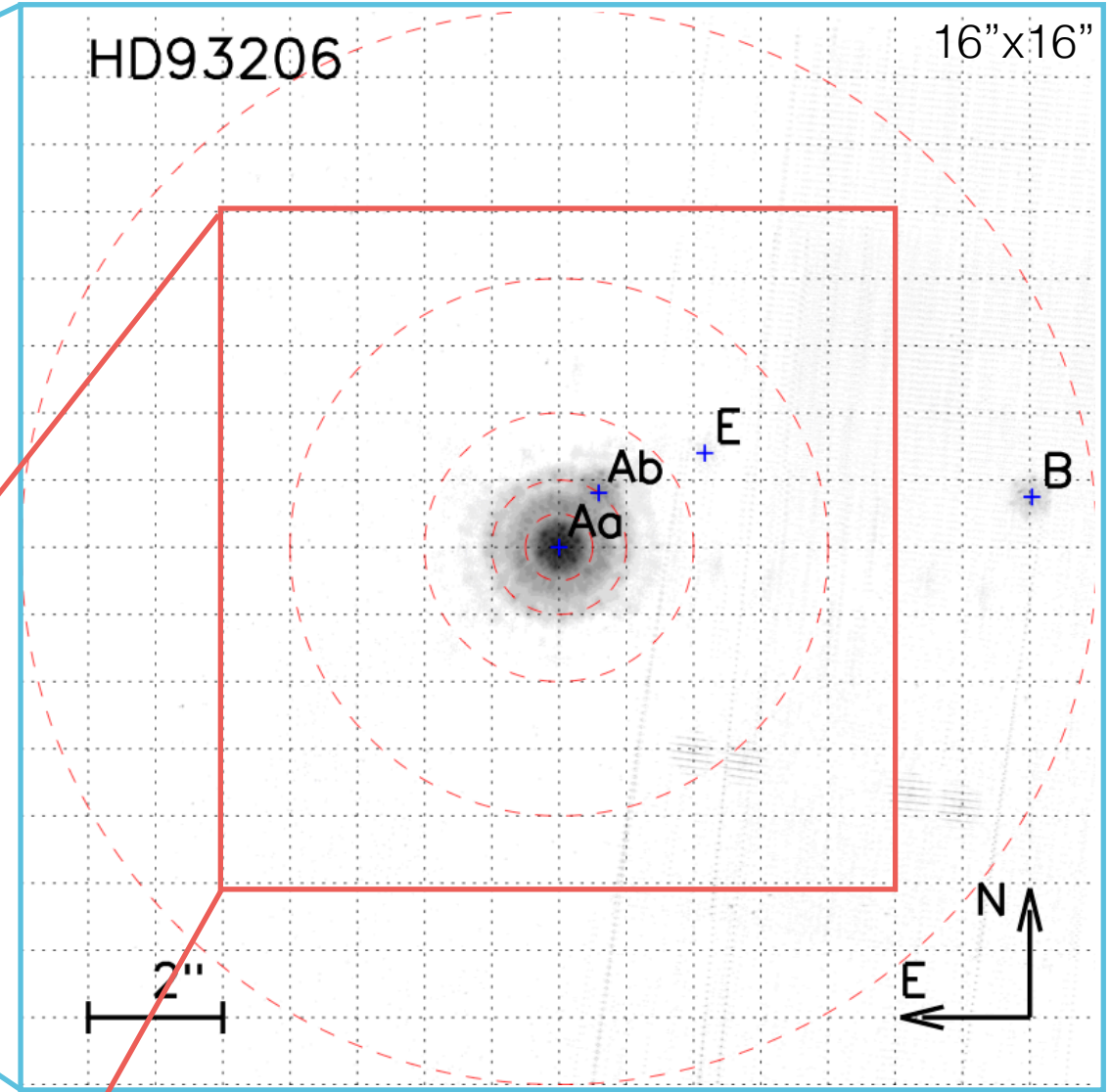


SMA+ Image  
Sana et al, 2014





FOV = 5', 2MASS



HD93206

16"x16"

Ab

E

Aα

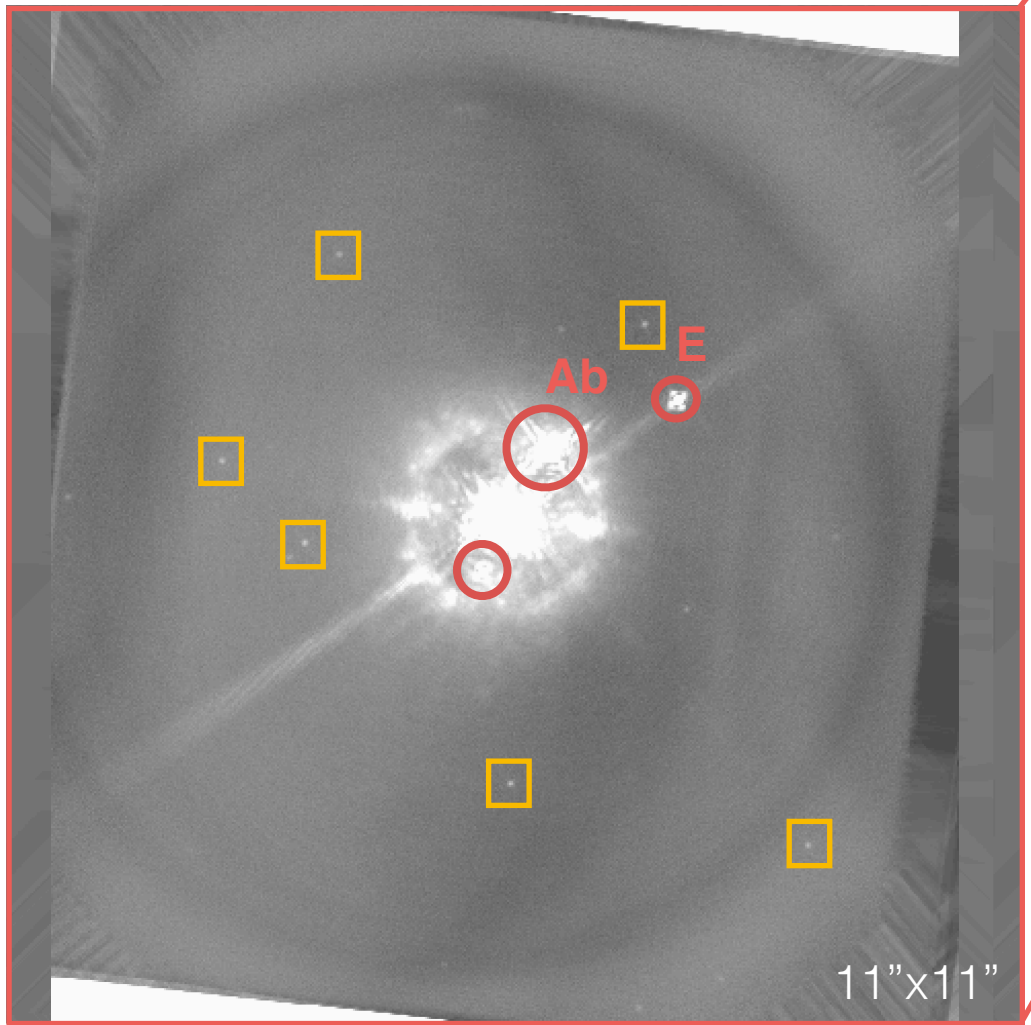
B

2"

N

E

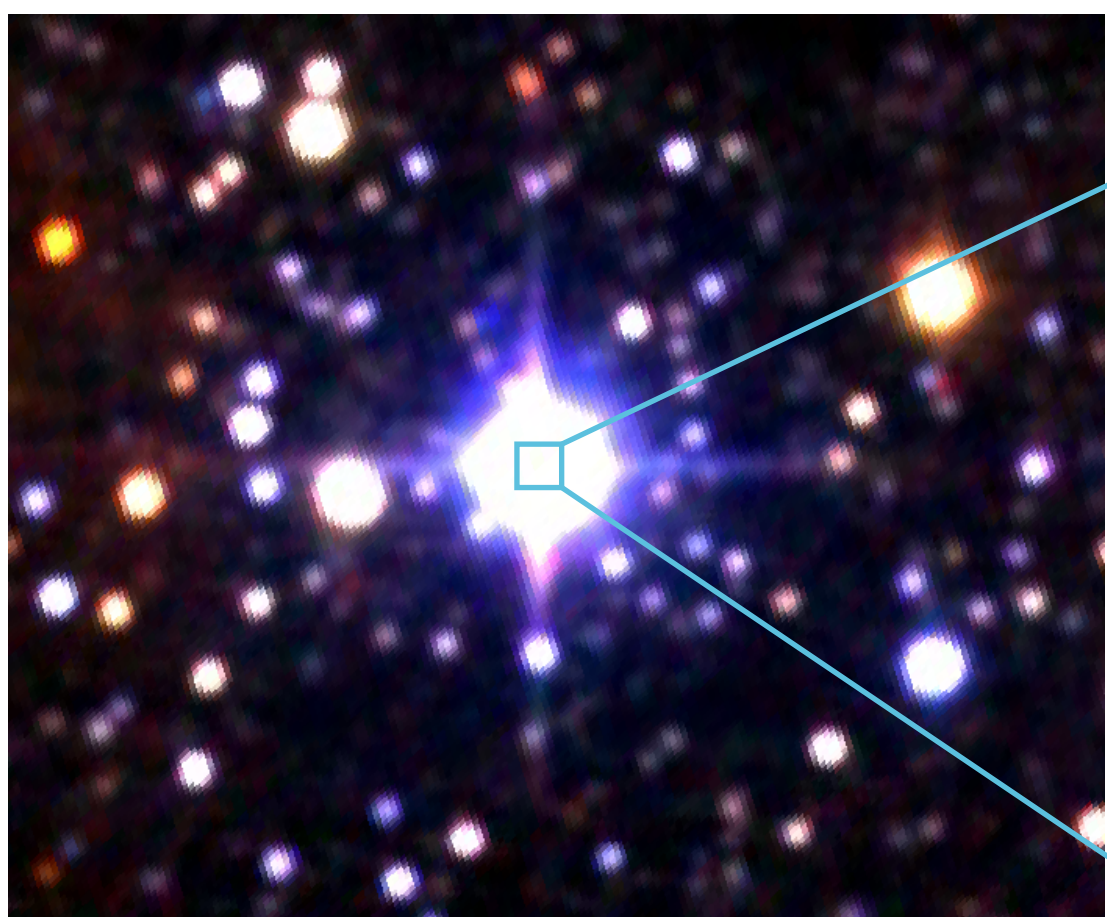
SMA+ Image  
Sana et al, 2014



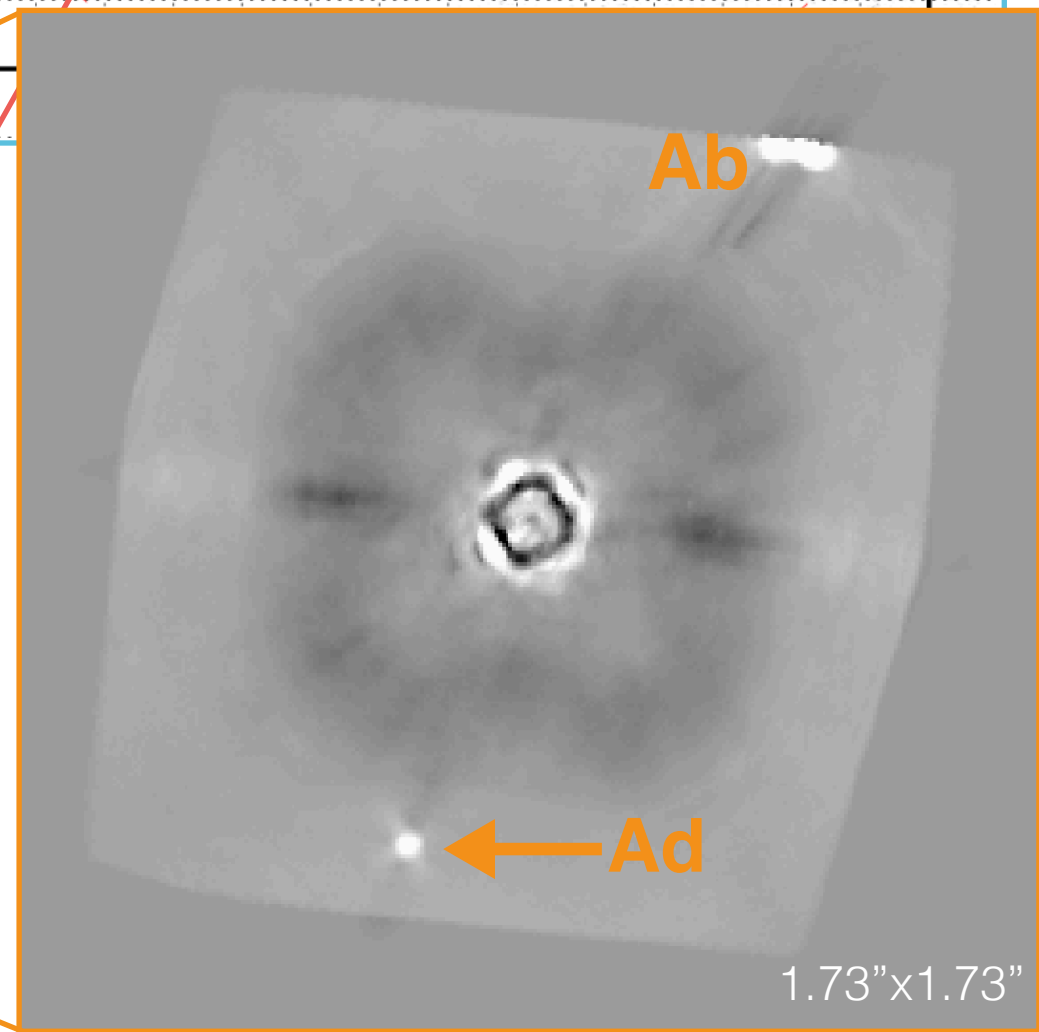
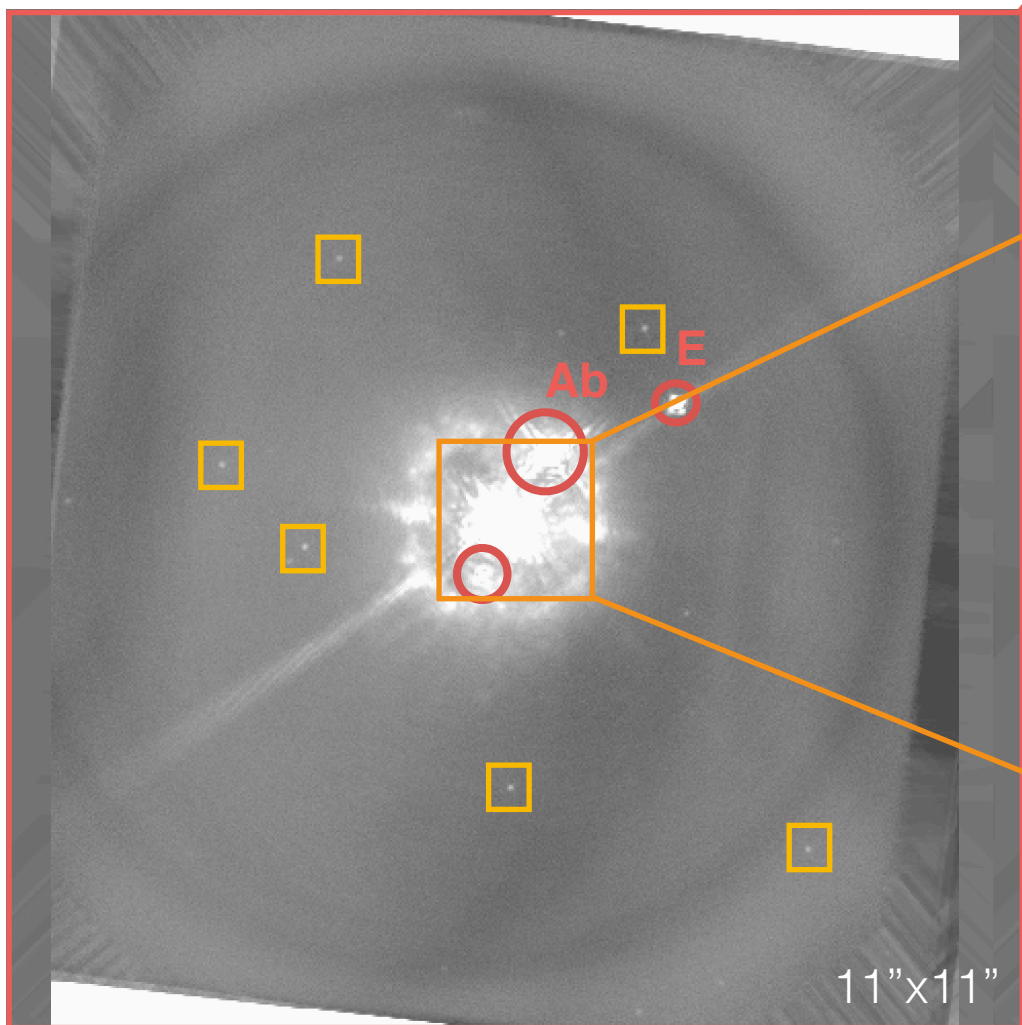
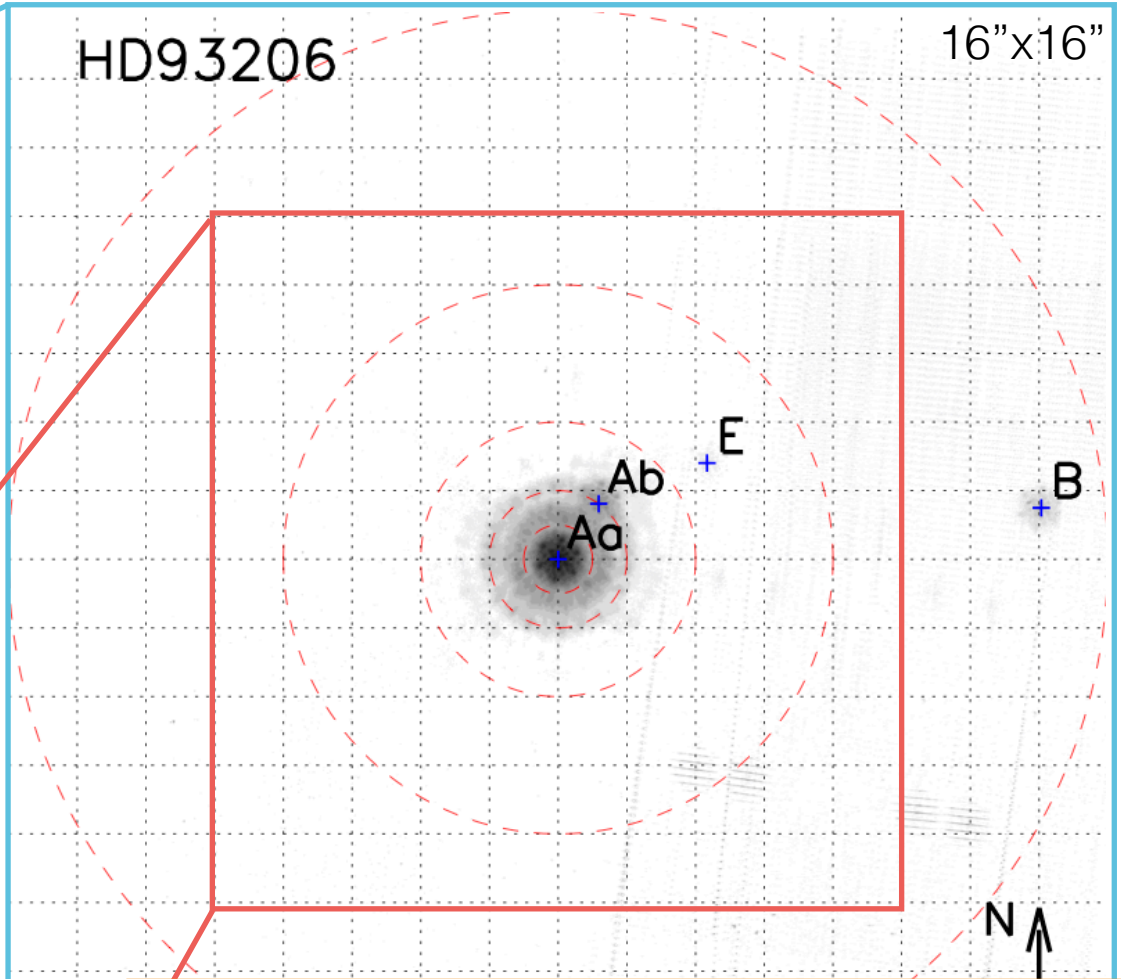
IRDIS

11"x11"



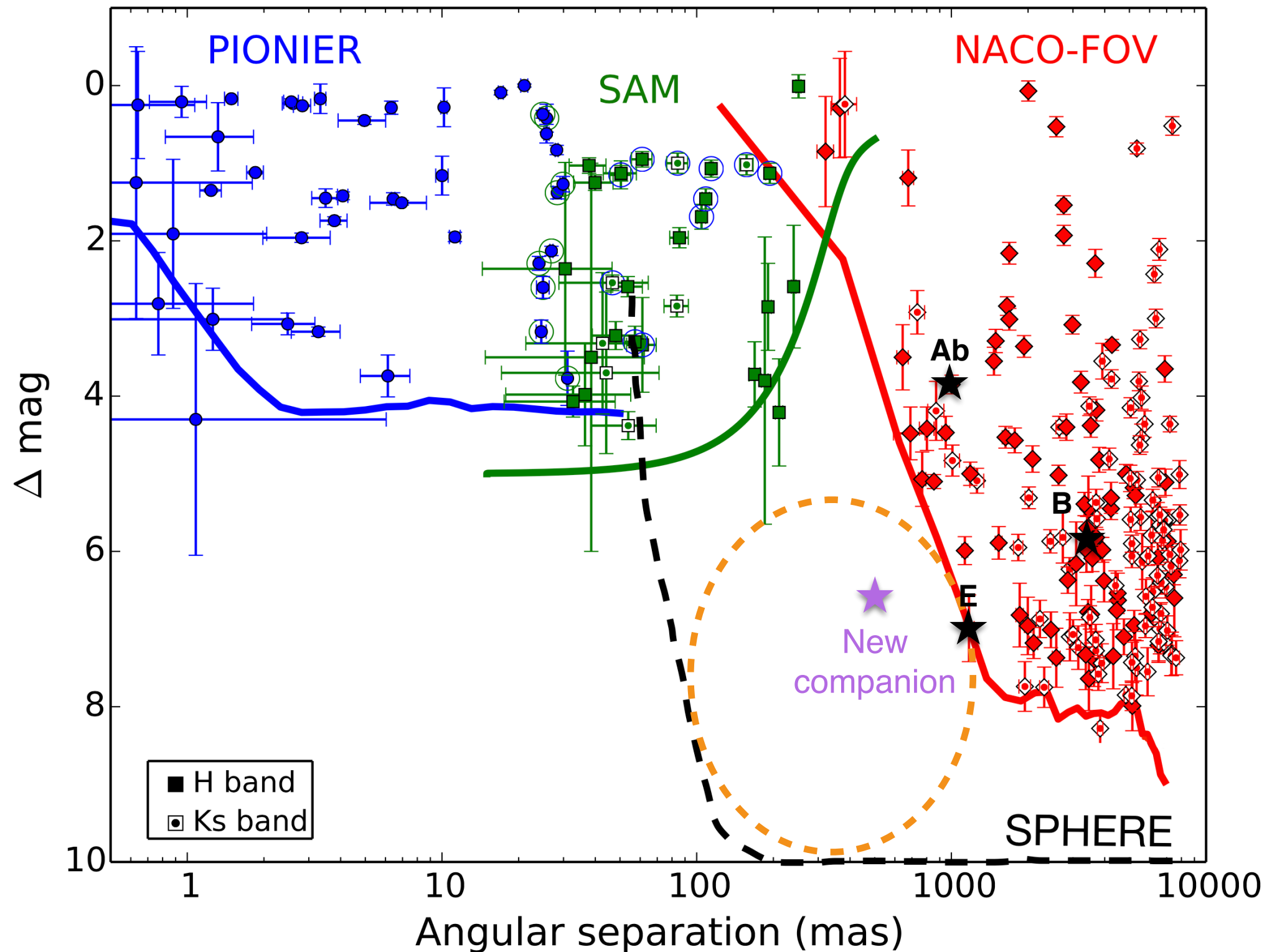


FOV = 5', 2MASS



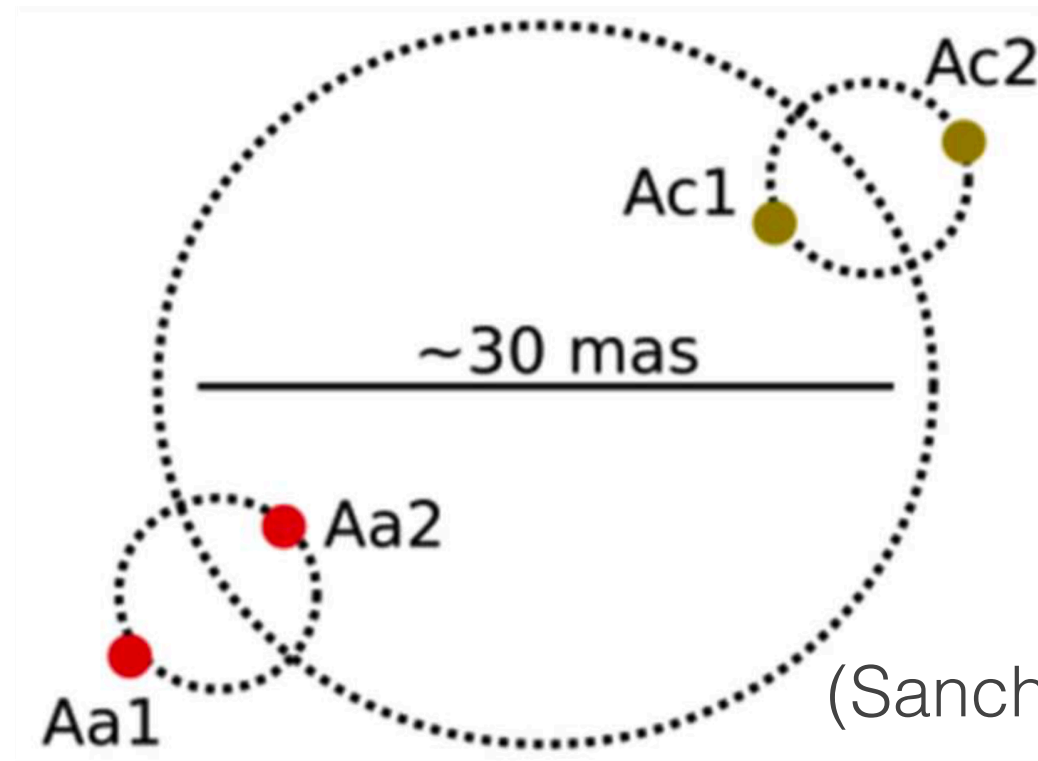


# SMaSH+ (Sana et al, 2014)



# Spectrum Extraction

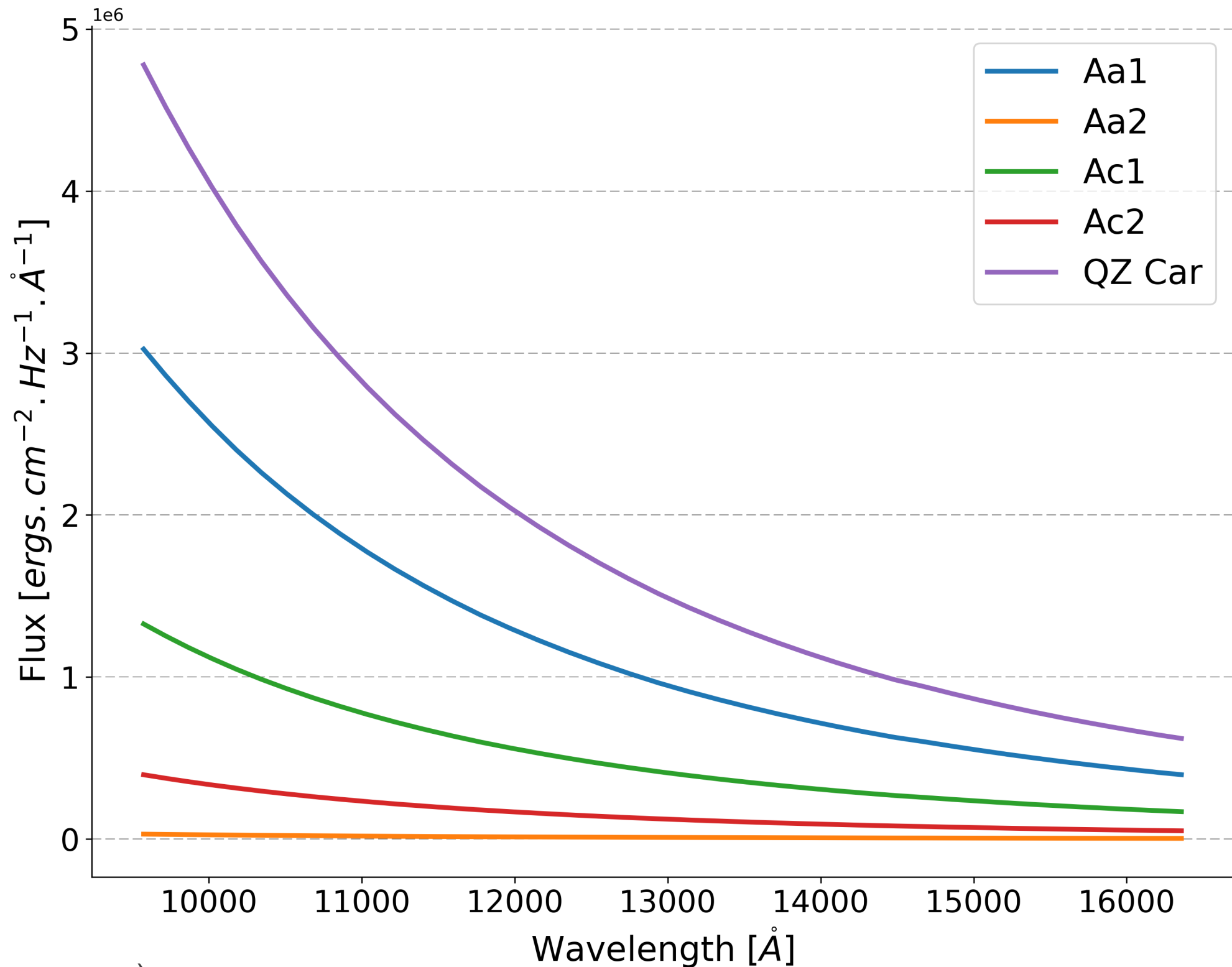
# QZ Car (HD 93206)



Component	Spectral Type	$T_{eff}$ (K)	$R_*$ ( $R_\odot$ )	$M_*$ ( $M_\odot$ )	$L_*$ ( $\log[L_*/L_\odot]$ )	$\dot{M}$ ( $M_\odot \cdot yr^{-1}$ )	$\log(g)$	$v_\infty$ ( $km \cdot s^{-1}$ )
Aa1	O9.7 I	32000	22.5	40	5.7	$8.21 \times 10^{-6}$	3.19	1794.3
Aa2	b2 v	20000	6.0	10	3.7	$2.39 \times 10^{-14}$	4.3	1186.4
Ac1	O8 III	32573	26.9	14.1	5.3	$3.32 \times 10^{-6}$	3.57	2191.2
Ac2	o9 v	32463	8.9	28	4.9	$3.16 \times 10^{-9}$	3.92	2427.1

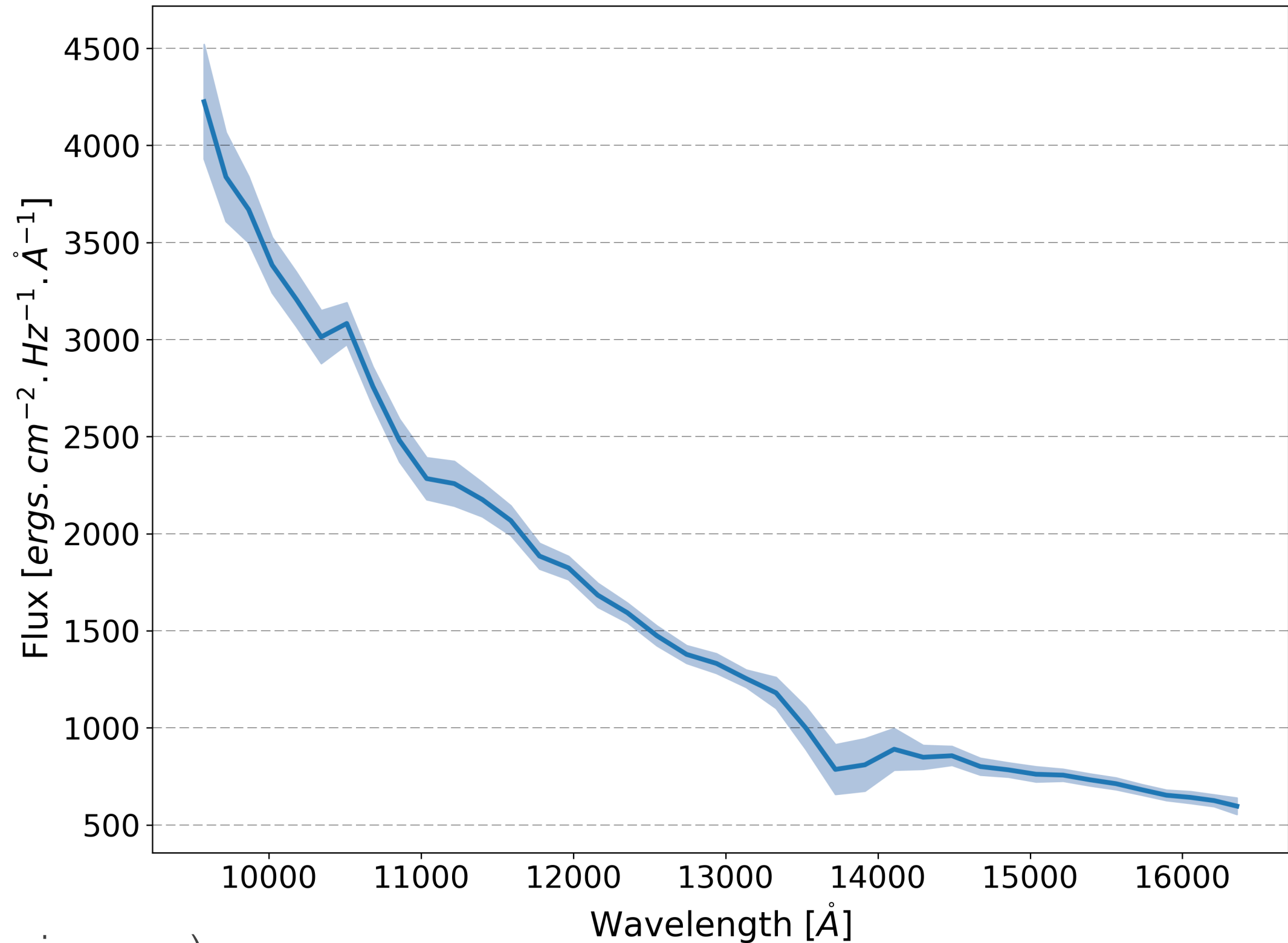
(Rainot+ in prep.)

# QZ Car spectrum with FASTWIND



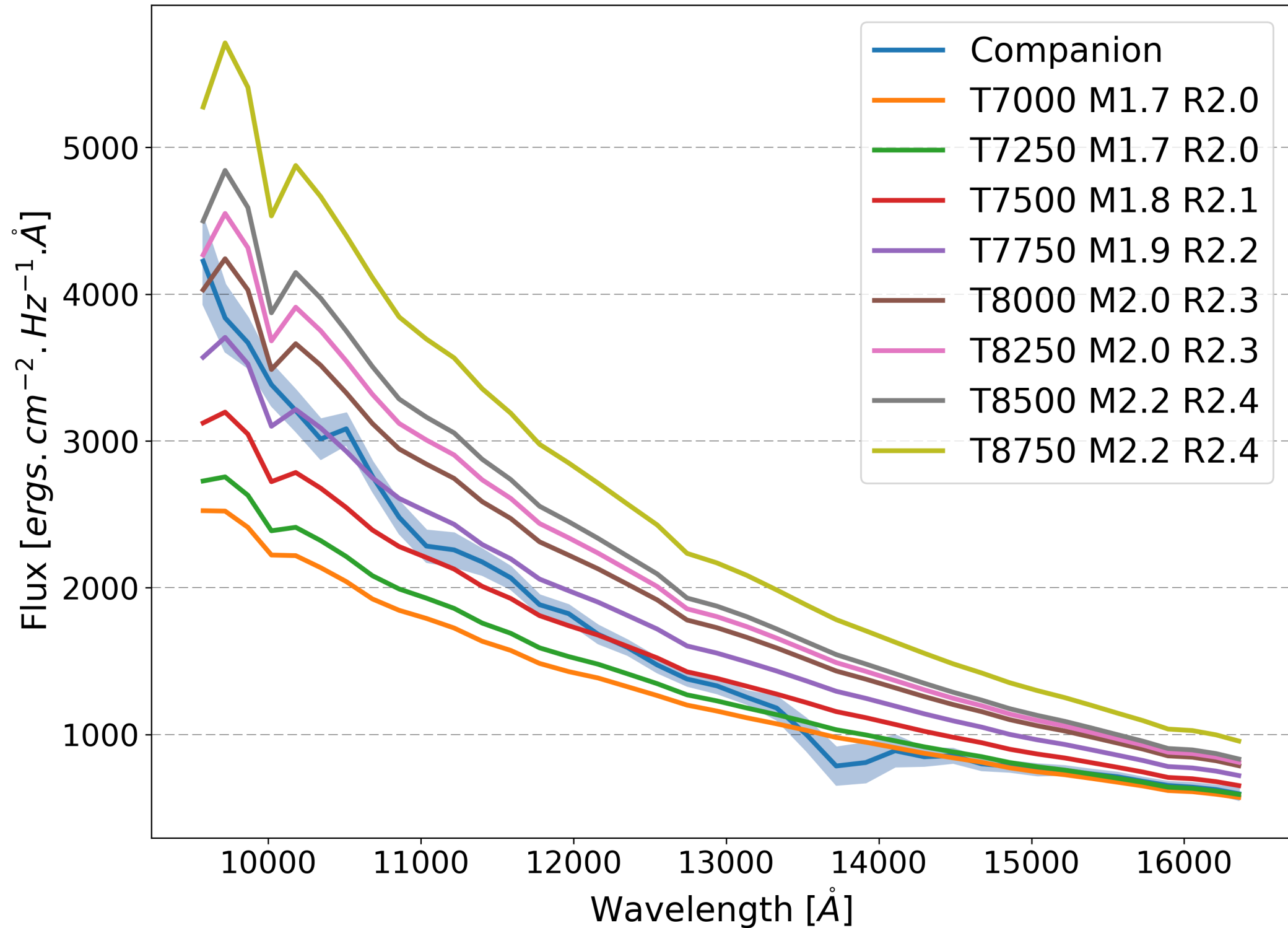
(Rainot+ in prep.)

# Calibrated spectrum



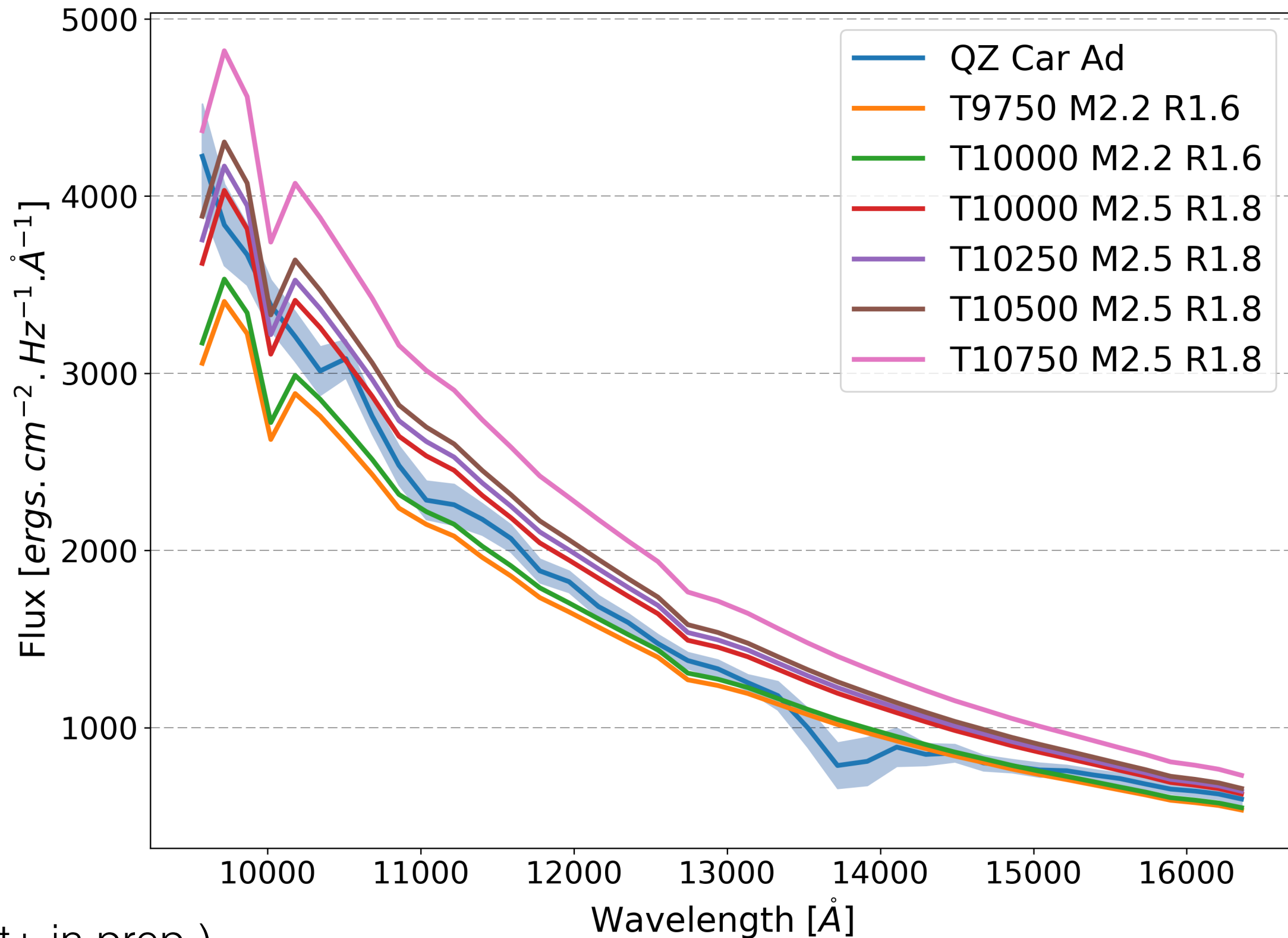
(Rainot+ in prep.)

# ZAMS



(Rainot+ in prep.)

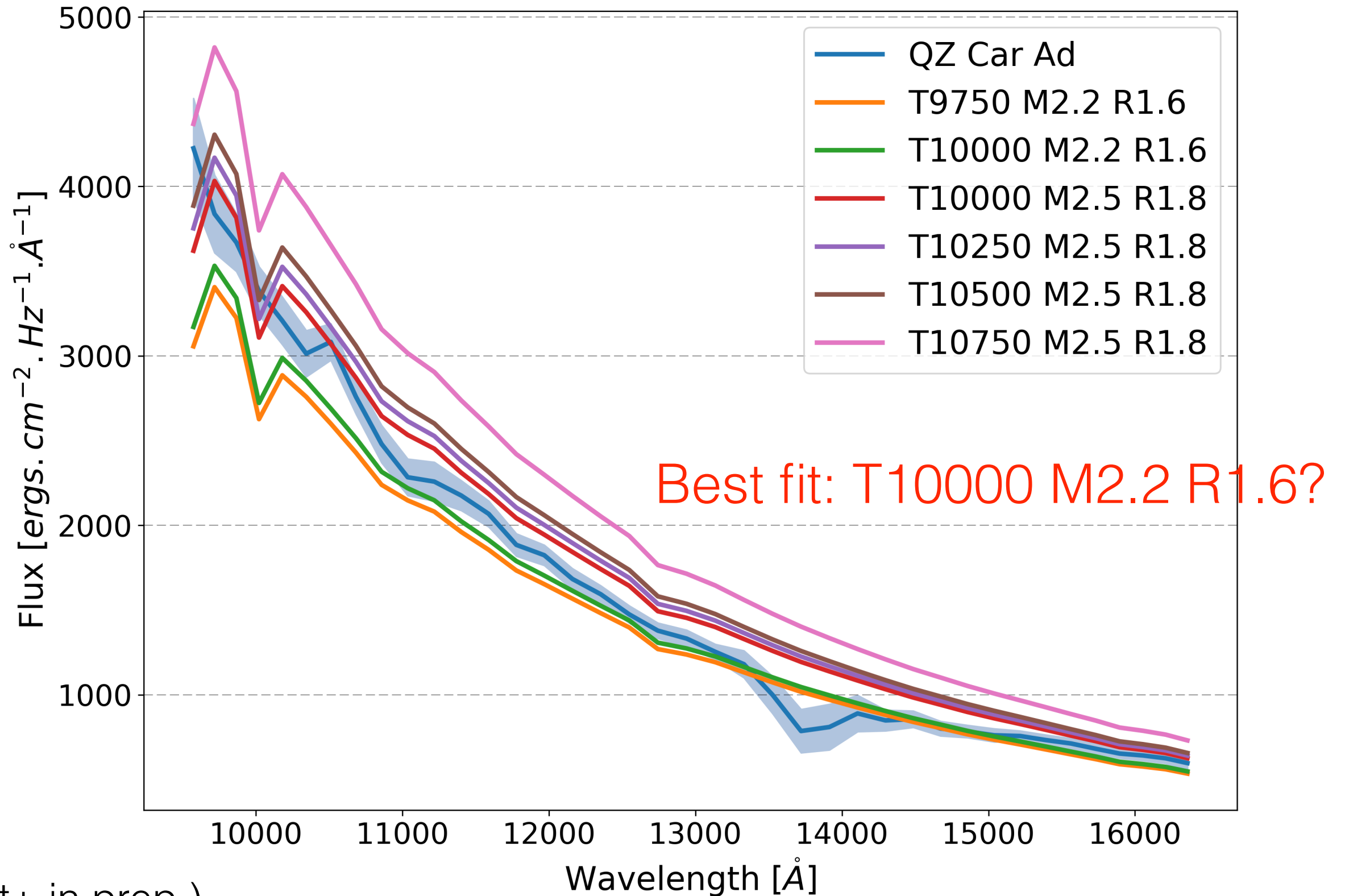
# Early MS



(Rainot+ in prep.)



# Early MS



(Rainot+ in prep.)

# Multiplicity Results - IFS

Images	Visible companions ( $\text{SNR} > 5\sigma$ )	Candidates ( $5\sigma > \text{SNR} > 3\sigma$ )
28	6	12

- ▶ Detection ratio  $\approx 0.42$  companions/star
- ▶ Expected  $\approx 38$  companions / 91 images

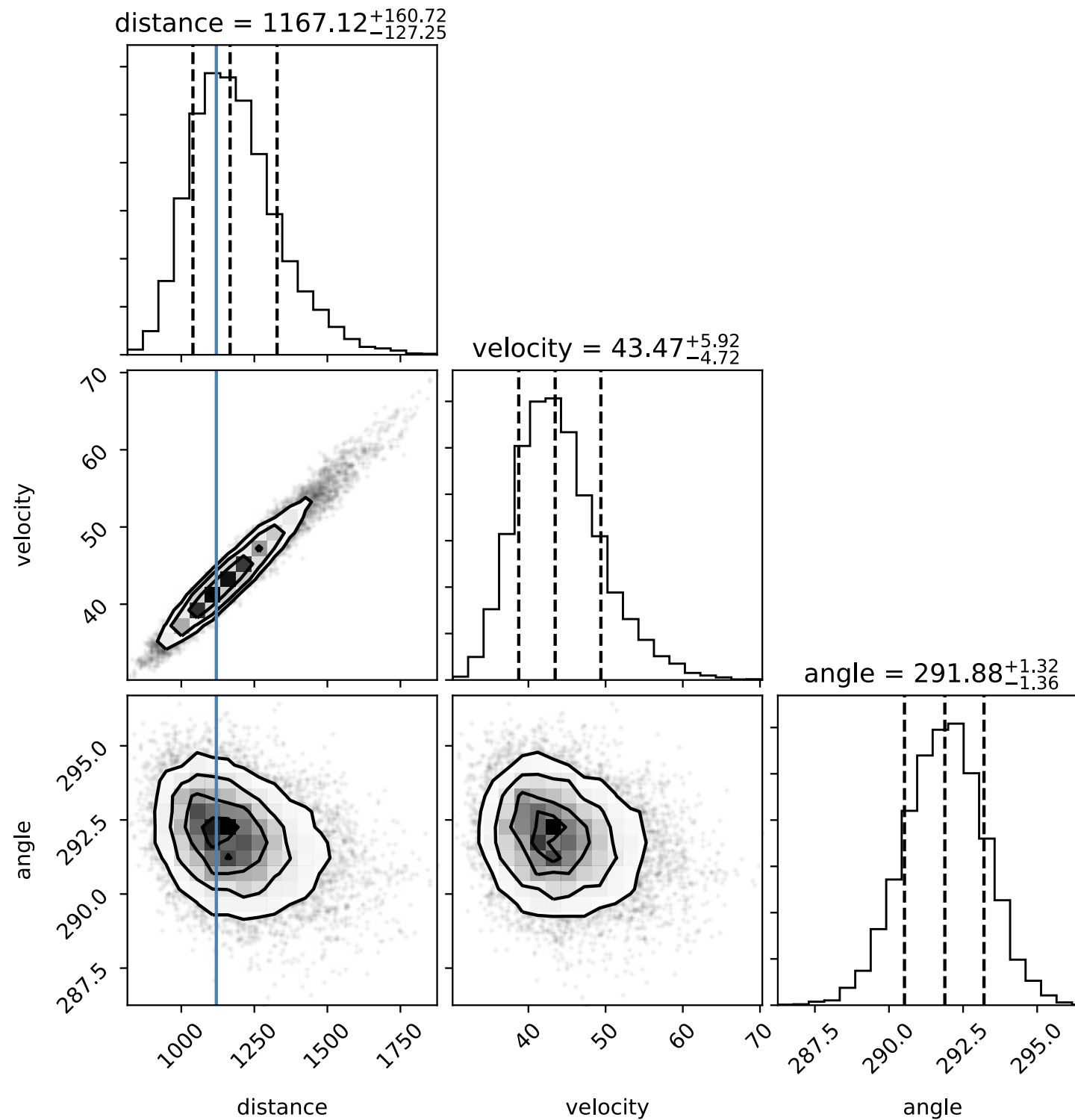
# Conclusion

- ▶ **Discovery** of a new companion: QZ Car Ad at 1.2kpc
- ▶ Might be an **A0 star** with:  $T = 10000\text{K}$ ,  $M = 2.2M$ ,  $R = 1.6R$
- ▶ Masses limits of IRDIS **detected sources**
- ▶ **63 stars** remain to be observed with SPHERE (P102)
- ▶ First paper to be submitted **soon!**

Thank you!

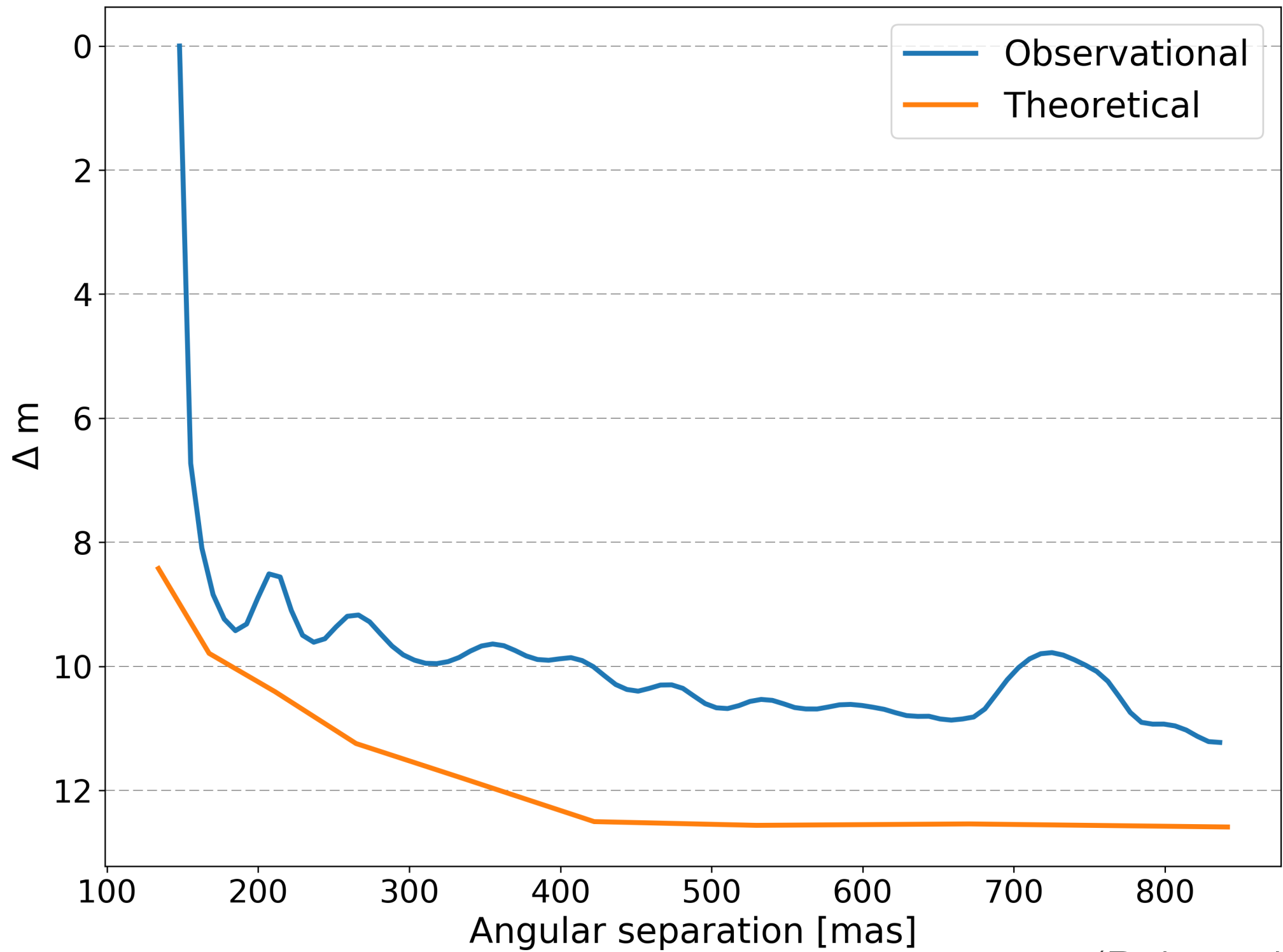
# Extra Slides

# GAI A distance



(Rainot+ in prep.)

# IFS Contrast Curves

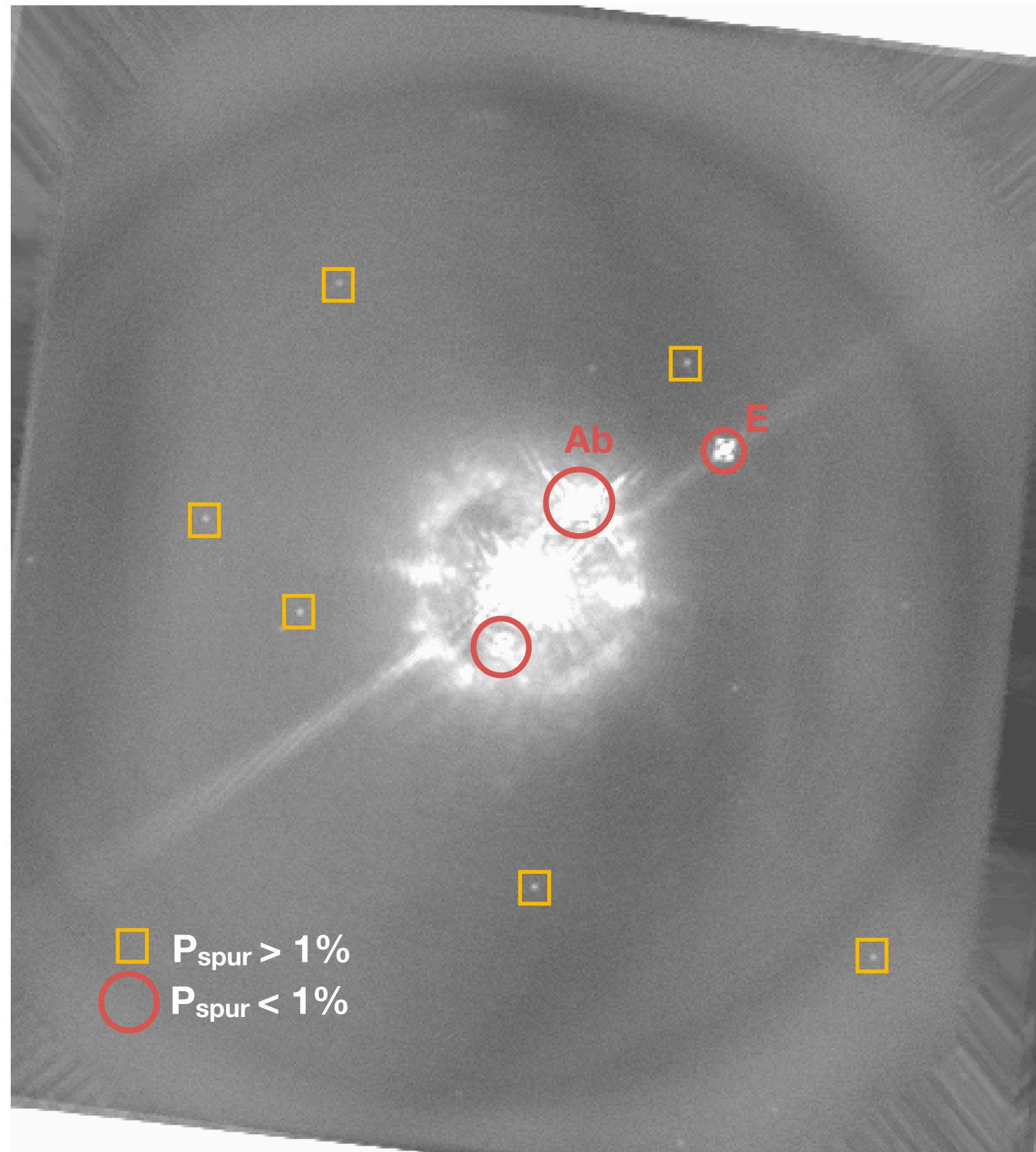


(Rainot+ in prep.)

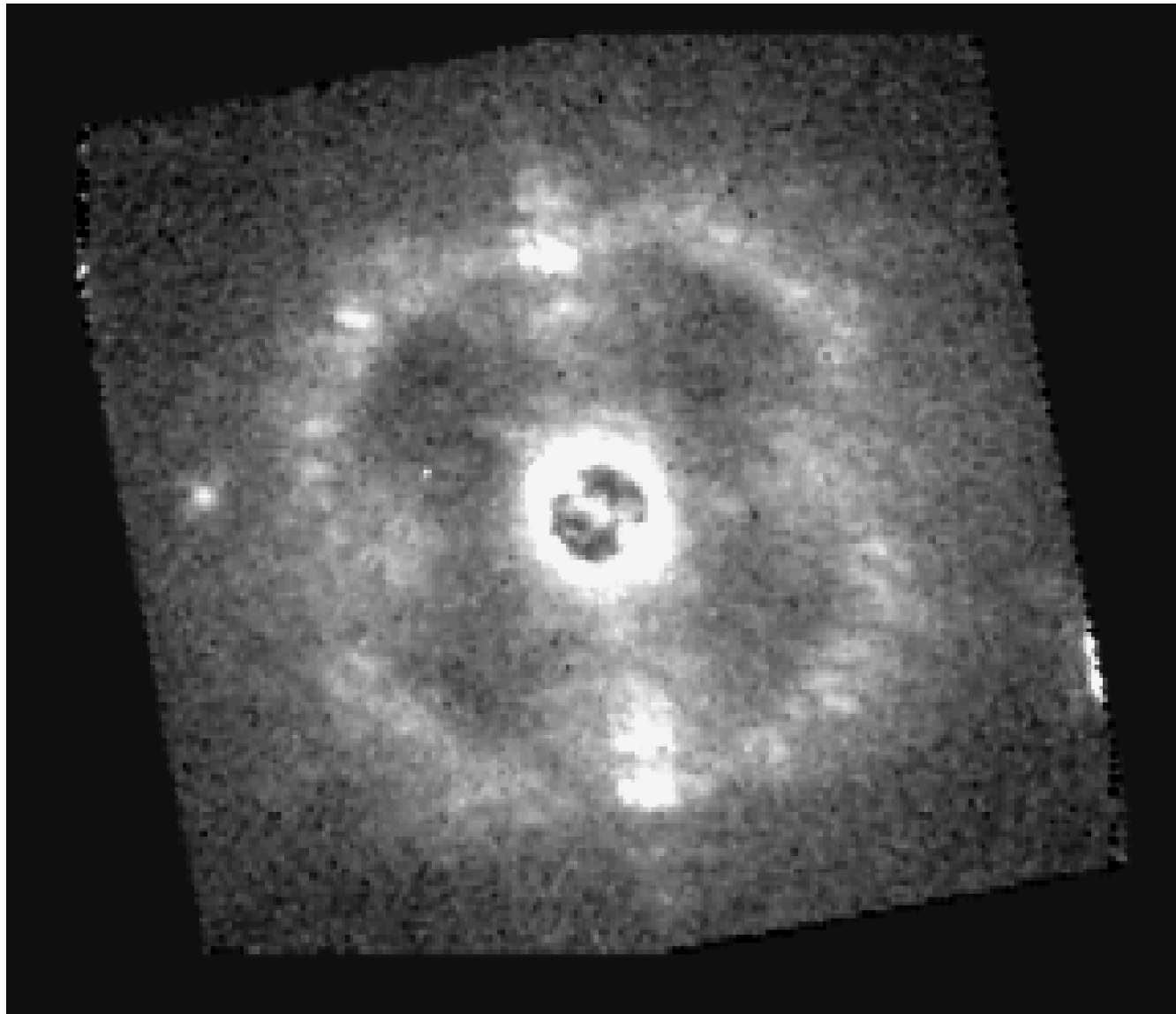


# IRDIS

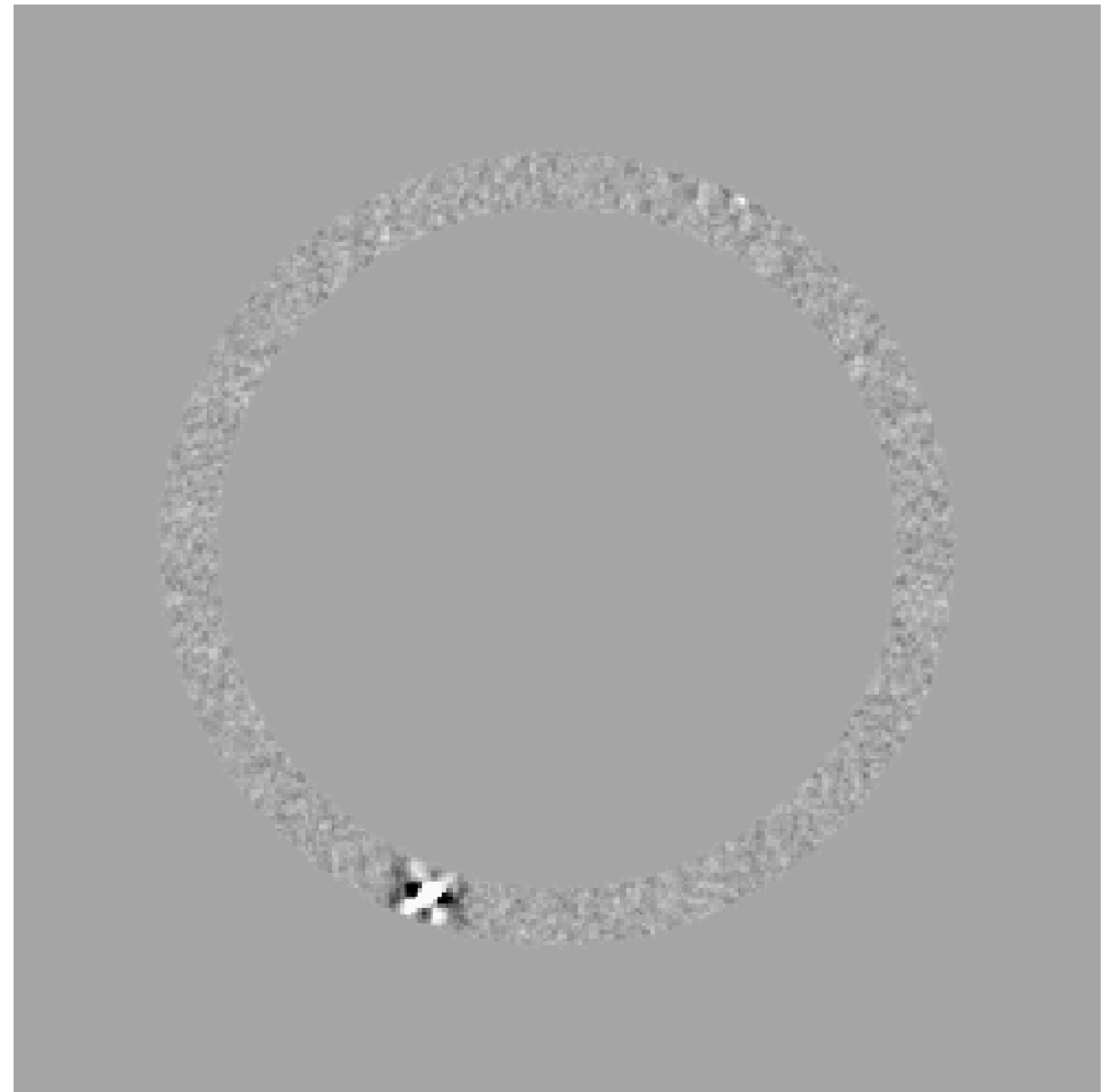
(Rainot+ in prep.)



# Analysis Techniques

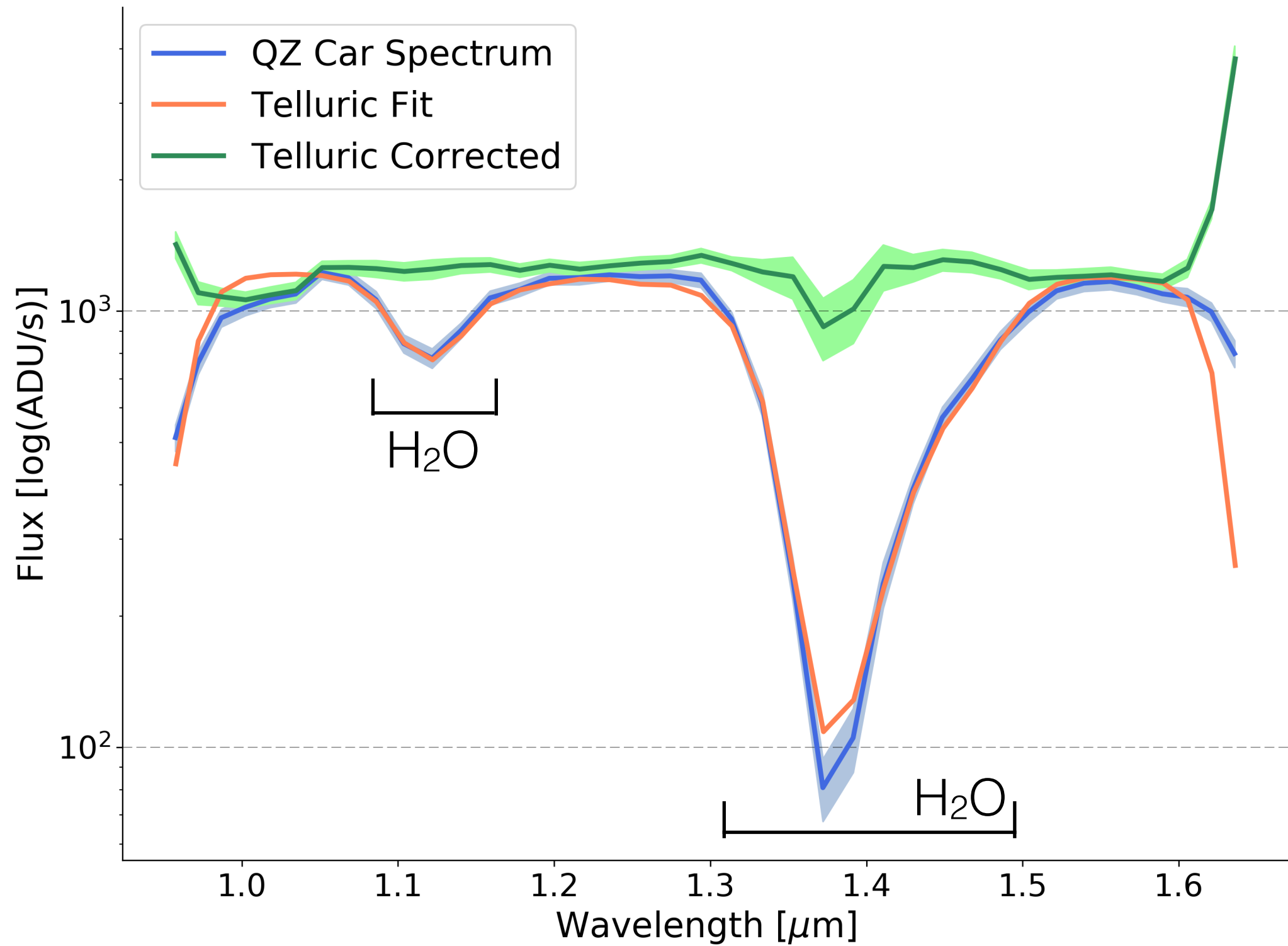


Original image

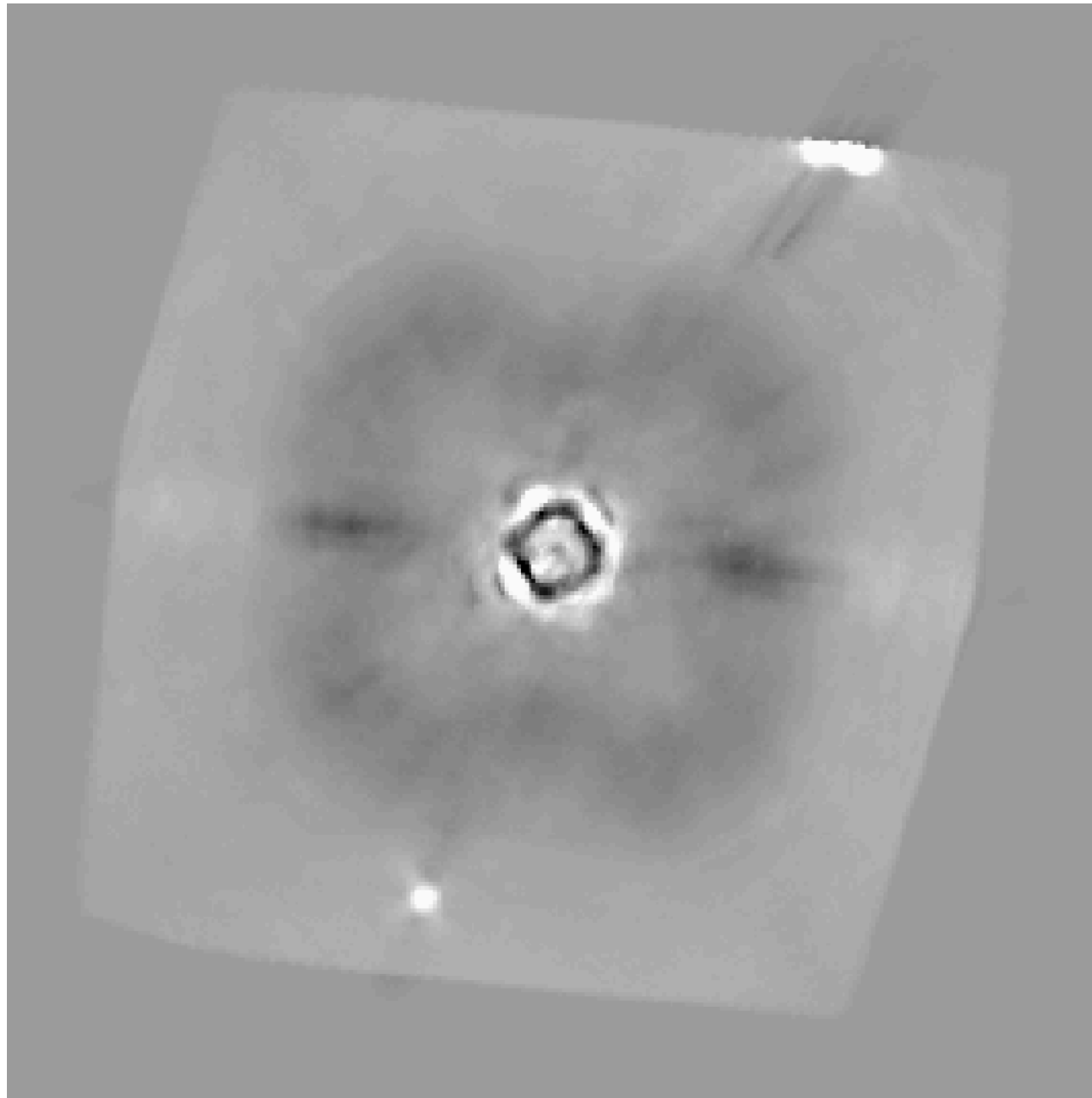


Post-processed cube

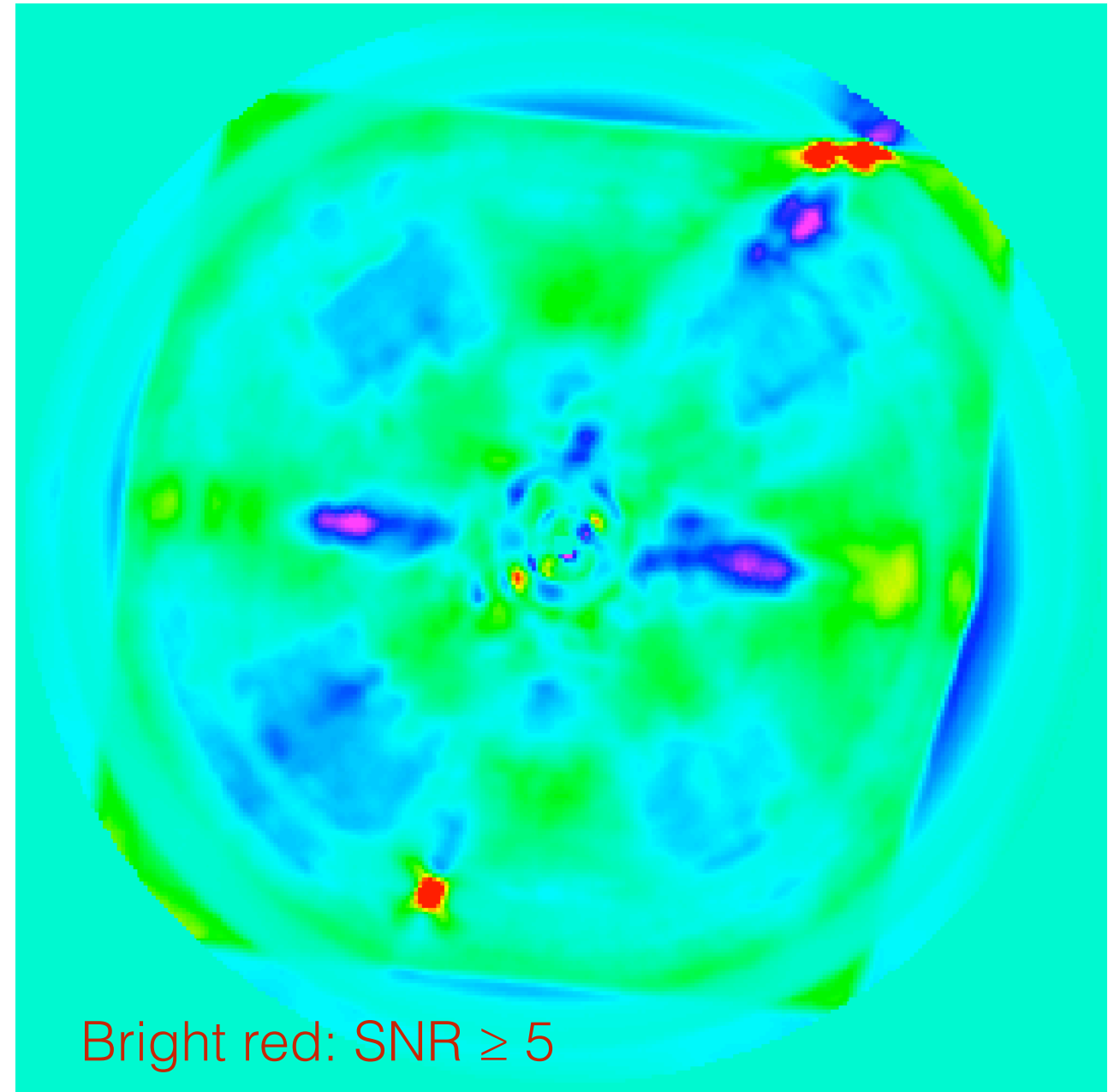
# PSF Fitting



# Detection method

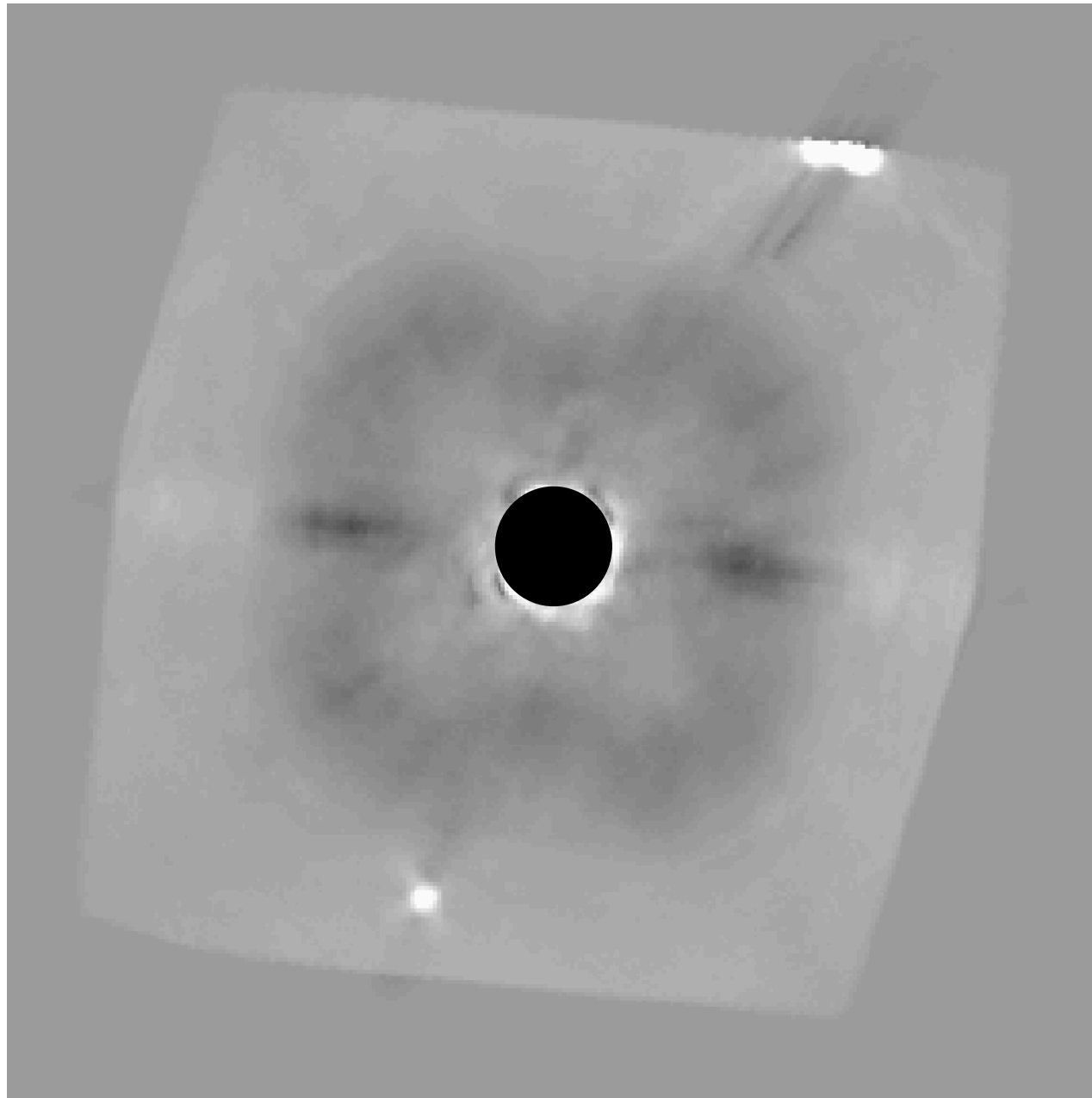


Derotated and wavelength collapsed image

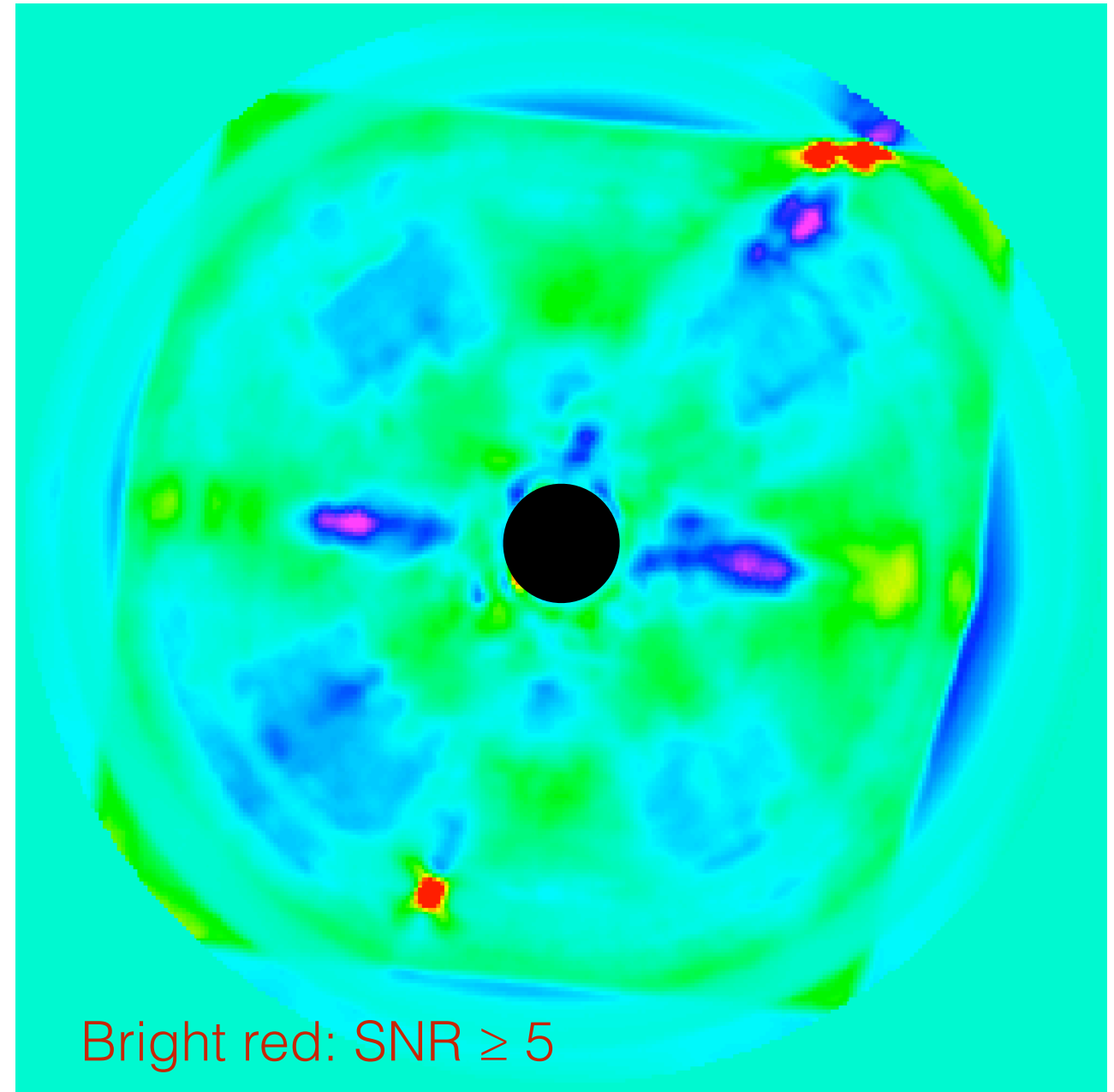


SNR map

# Detection method



Derotated and wavelength collapsed image



SNR map